

Problem sletanja na južnu hemisferu Marsa

Dušan Marčeta

Katedra za Astronomiju

Matematički fakultet

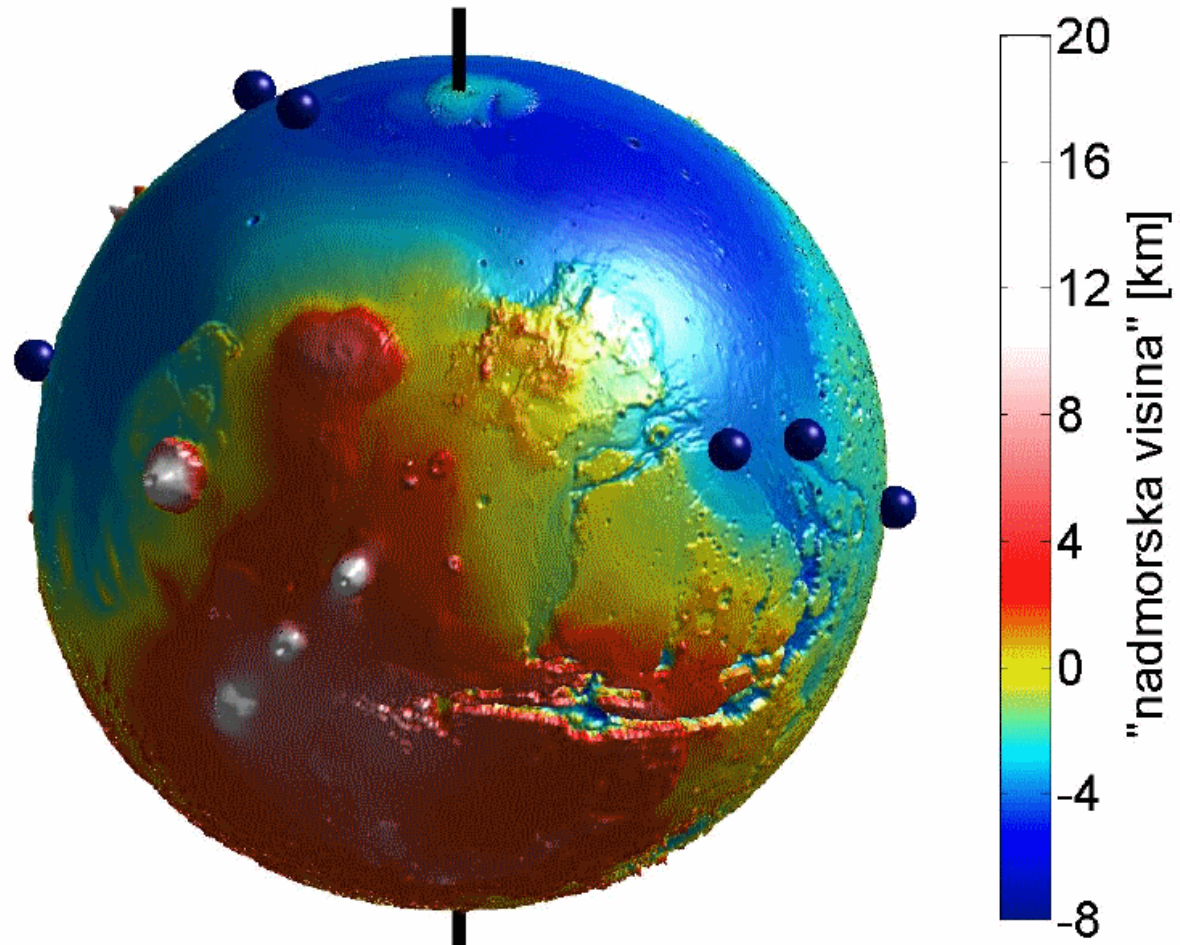
Univerzitet u Beogradu

Departman za fiziku, PMF Novi Sad,
06.11.2015.

CILJEVI ISTRAŽIVANJA

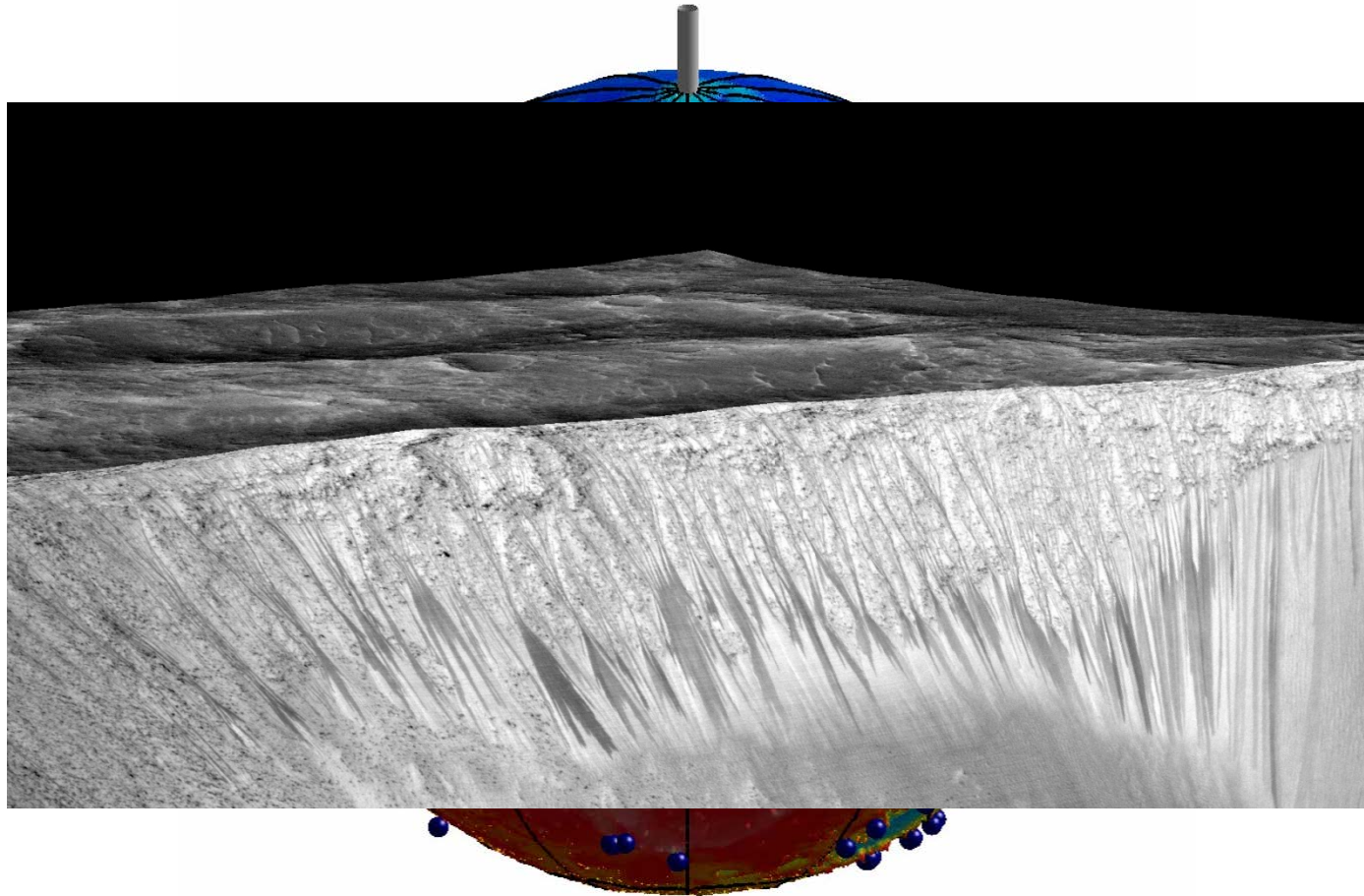
- Razvoj algoritma za određivanje balističke putanje objekta kroz planetsku atmosferu
- Istraživanje mogućnosti za sletanje na južne visoravni, bez unapređivanja postojeće tehnologije
- Određivanje optimalnih parametra za sletanje (L_s , $LPSV$, v , γ)

MOTIVACIJA



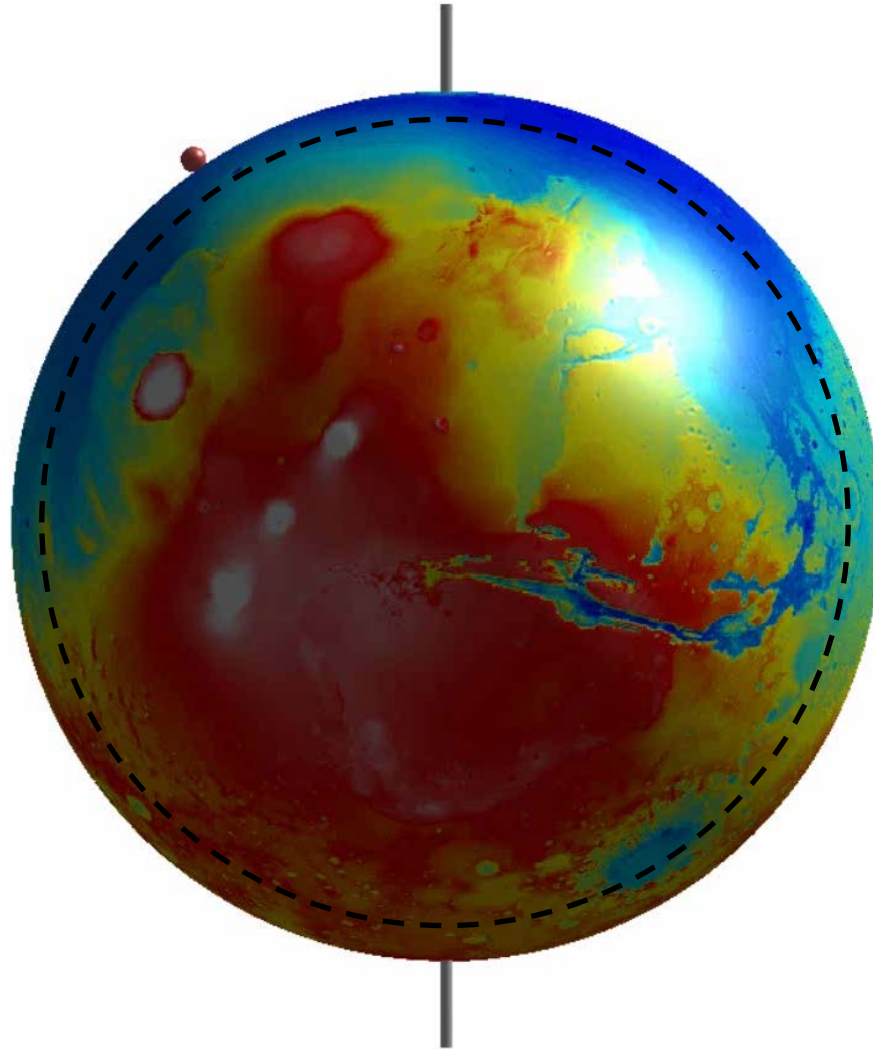
Smith et al., *Science*, 1999.

MOTIVACIJA



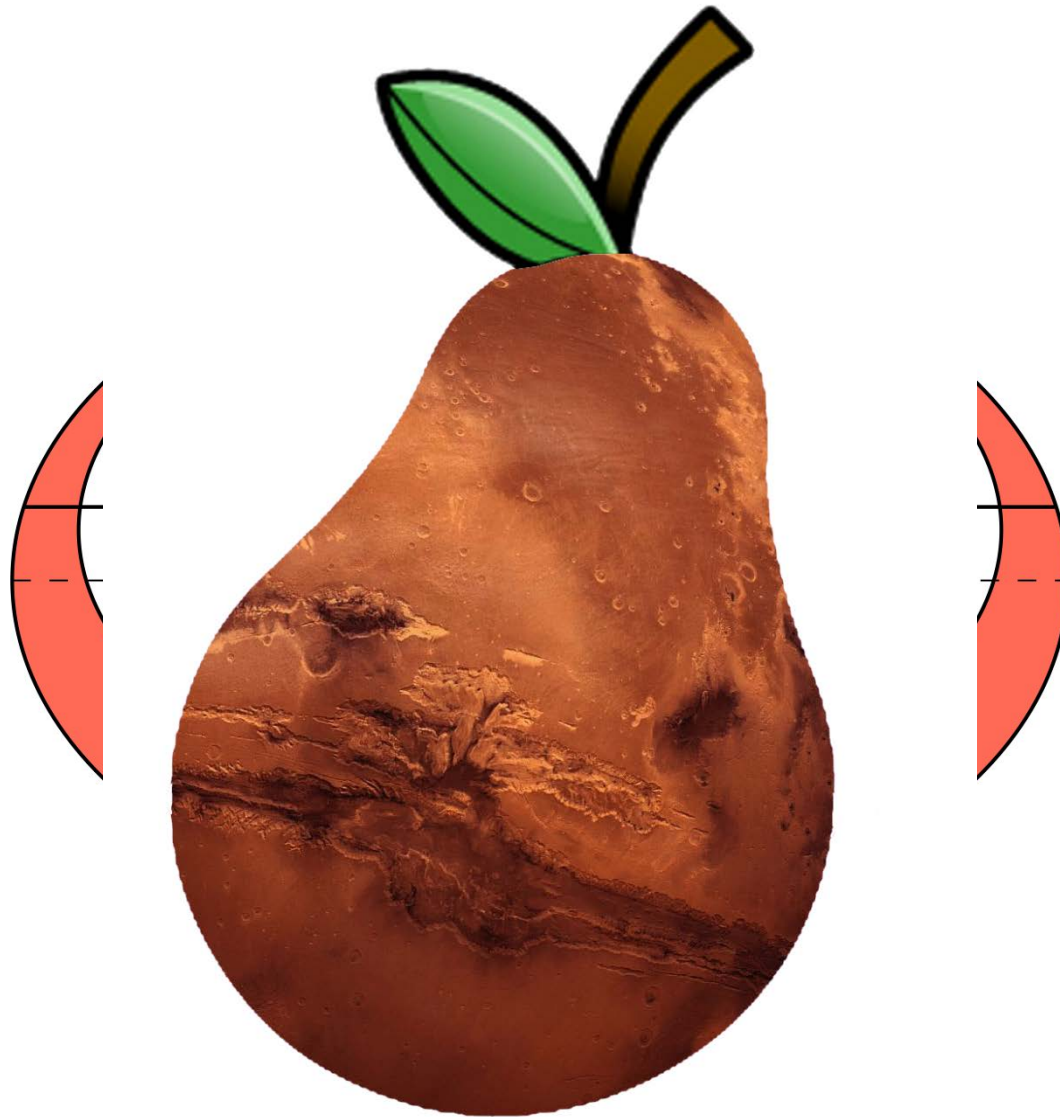
Fassett & Head, *Icarus*, 2008.

POREKLO GLOBALNE DIHOTOMIJE



Wilhelms & Squyres, *Nature*, 1984.

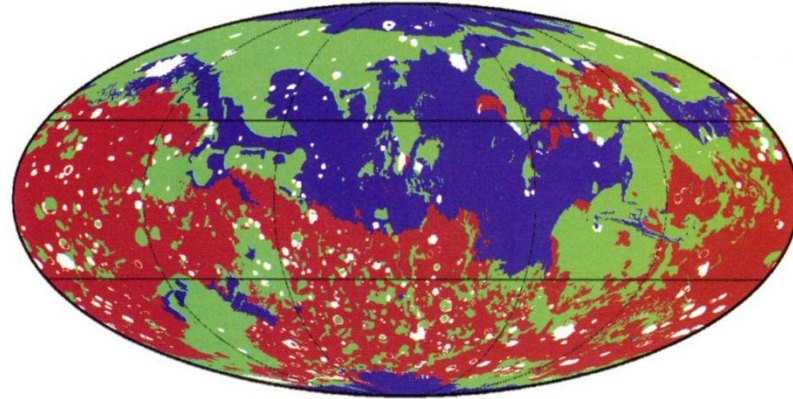
UNUTRAŠNJA STRUKTURA MARSA



Smith et al., *Science*, 1999.

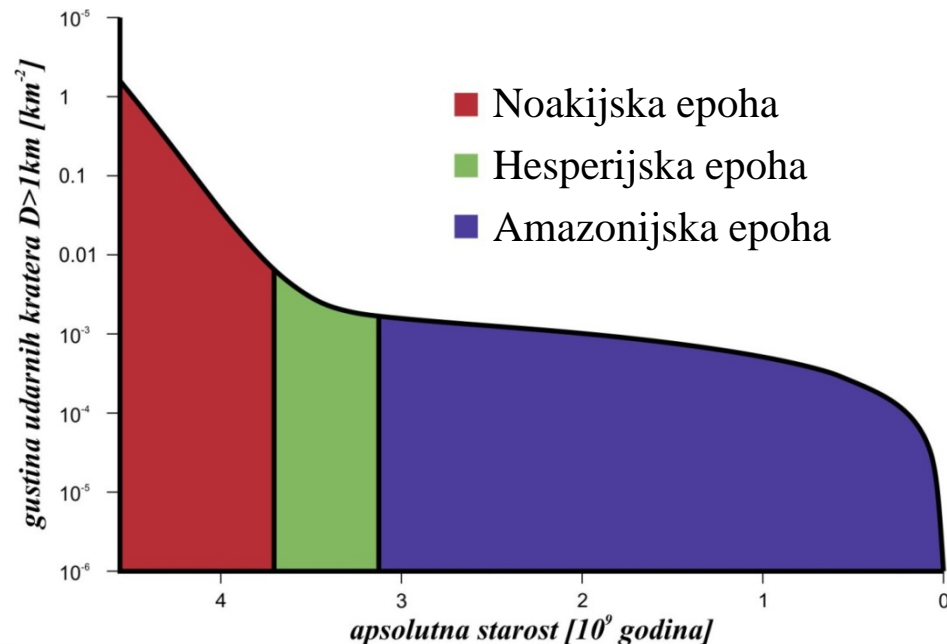
STRATIGRAFIJSKA PODELA POVRŠINE MARSA

Solomon et al., *Science*, 2005.

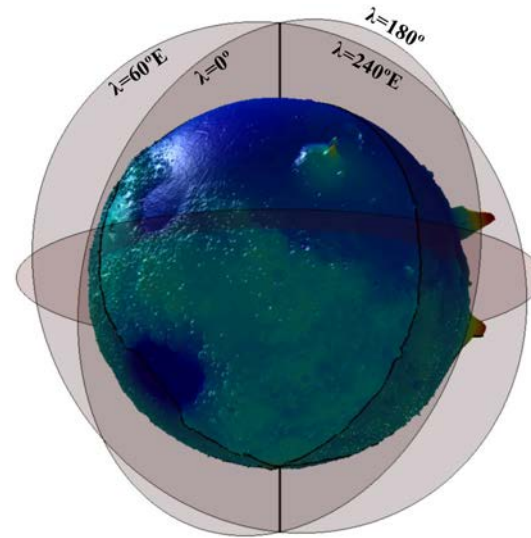
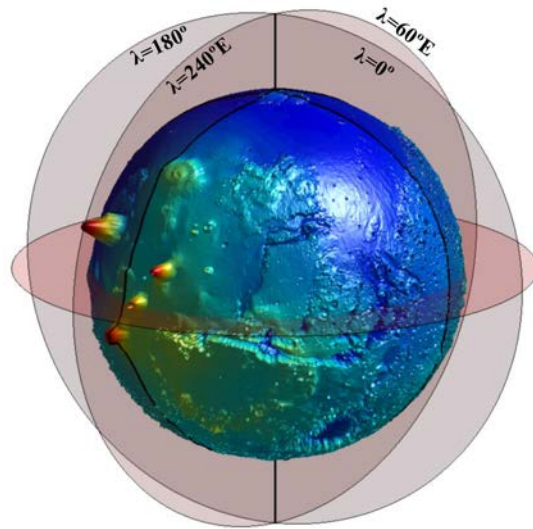


$$N(1km) = 2.68 \cdot 10^{-14} (e^{6.93t} - 1) + 4.131 \cdot 10^{-4}t$$

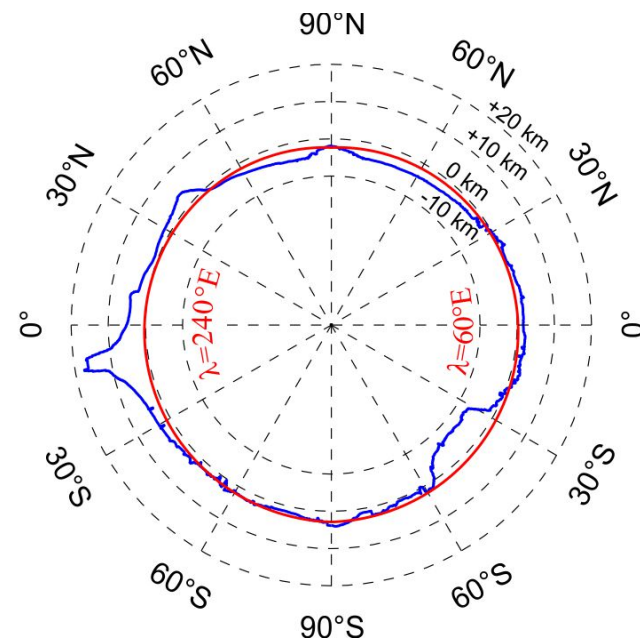
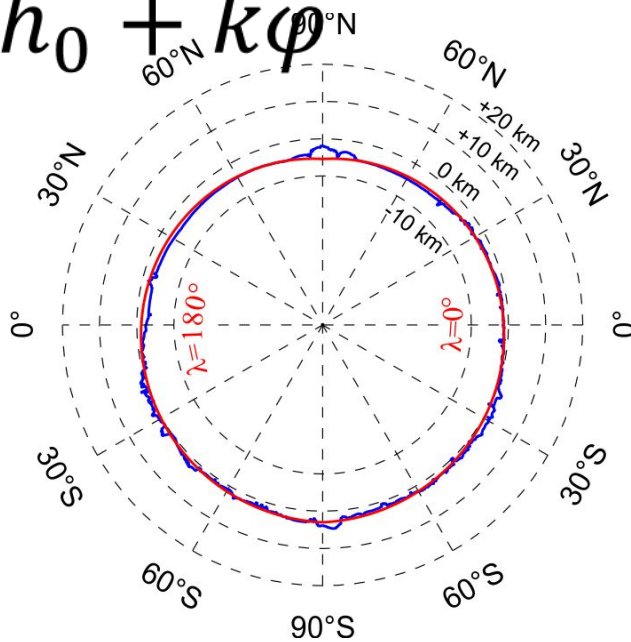
Neukum, *Lunar and Planetary Science Conference*, 2008



TOPOGRAFIJA MARSOVE POVRŠINE



$$h = h_0 + k\varphi$$



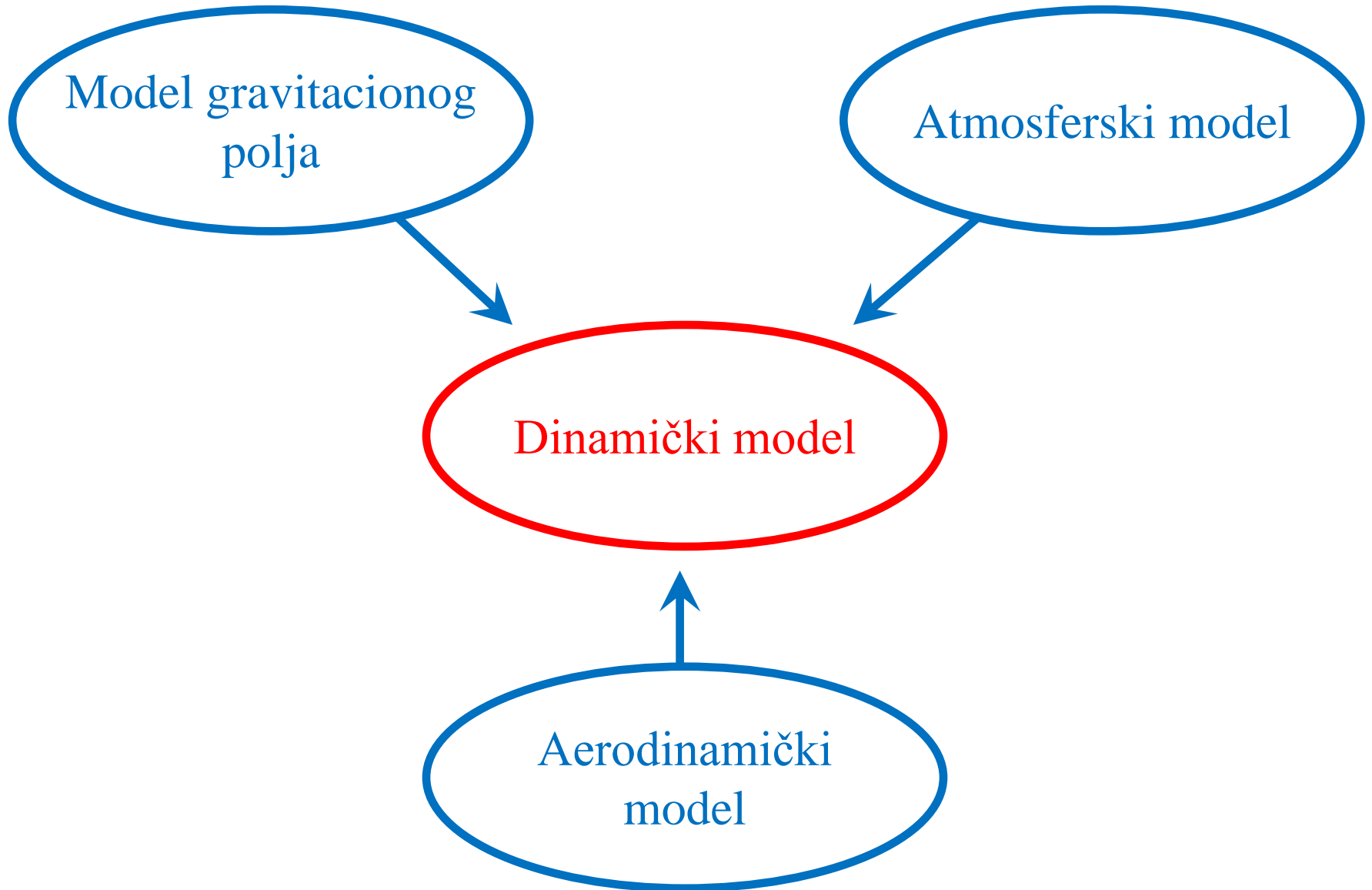
Mars vs Zemlja



R



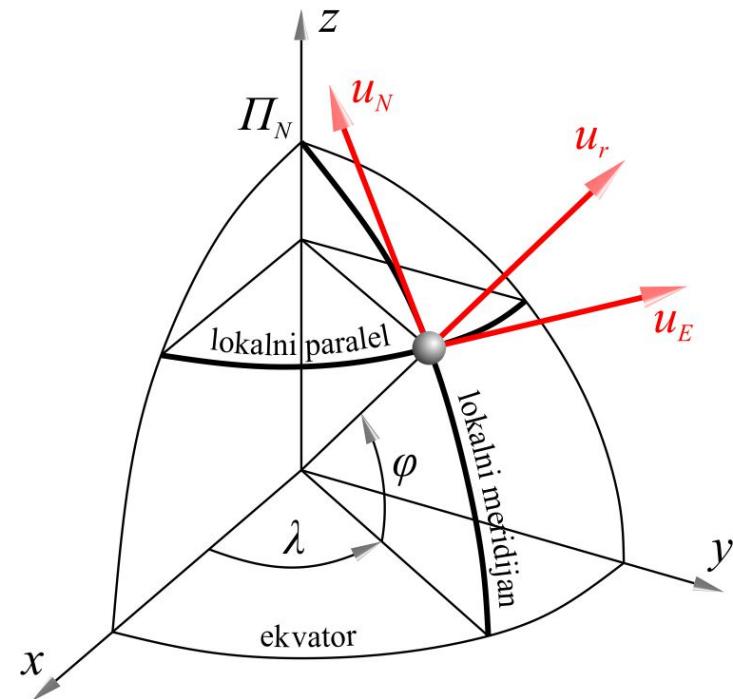
ATMOSFERSKA PUTANJA



GRAVITACIONO POLJE MARSA

$$U = \frac{GM}{r} \left\{ 1 + \sum_{l=2}^{\infty} \left(\frac{r_0}{r}\right)^l \sum_{m=0}^l P_{lm}(\sin \delta) [C_{lm} \cos m\lambda + S_{lm} \sin m\lambda] \right\}$$

$$\vec{a} = \frac{\partial U}{\partial r} \vec{u}_r + \frac{1}{r} \frac{\partial U}{\partial \delta} \vec{u}_N + \frac{1}{r \cos \delta} \frac{\partial U}{\partial \lambda} \vec{u}_E$$



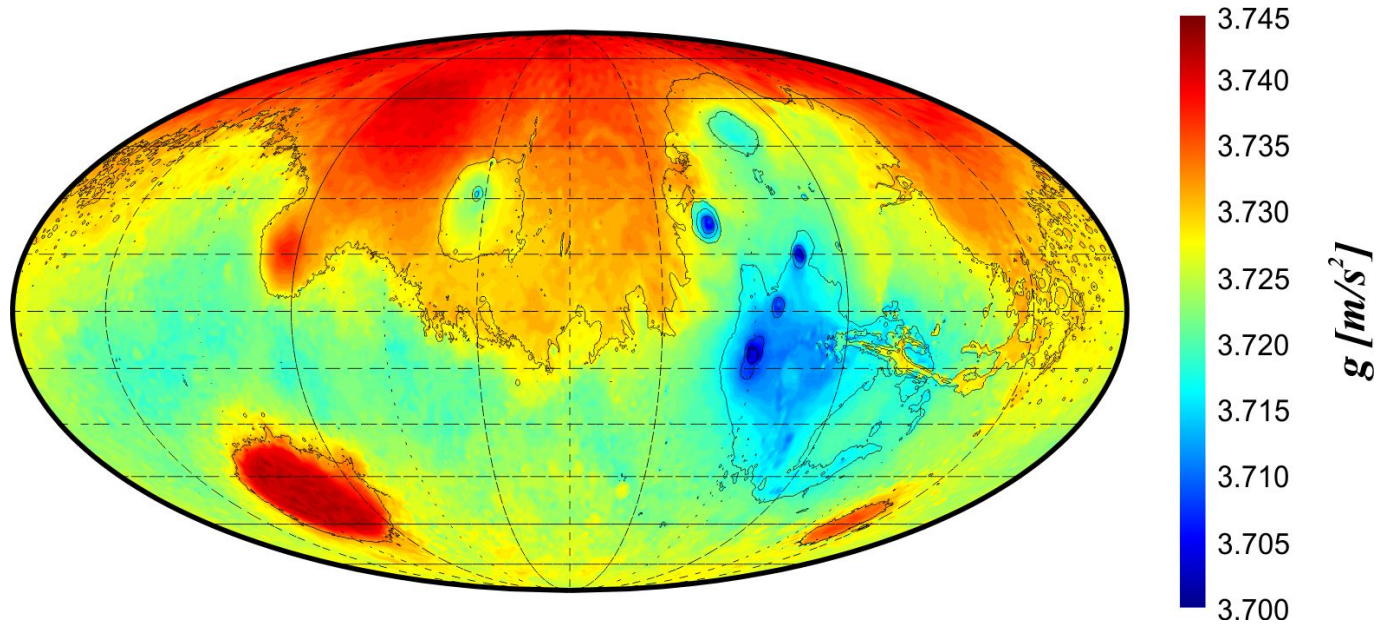
UBRZANJA GRAVITACIONOG POLJA

$$a_r = -\frac{GM}{r^2} \left(1 + \sum_{l=2}^{\infty} (l+1) \left(\frac{r_0}{r}\right)^l \sum_{m=0}^l P_{lm}(\sin \delta) [C_{lm} \cos m\lambda + S_{lm} \sin m\lambda] \right)$$

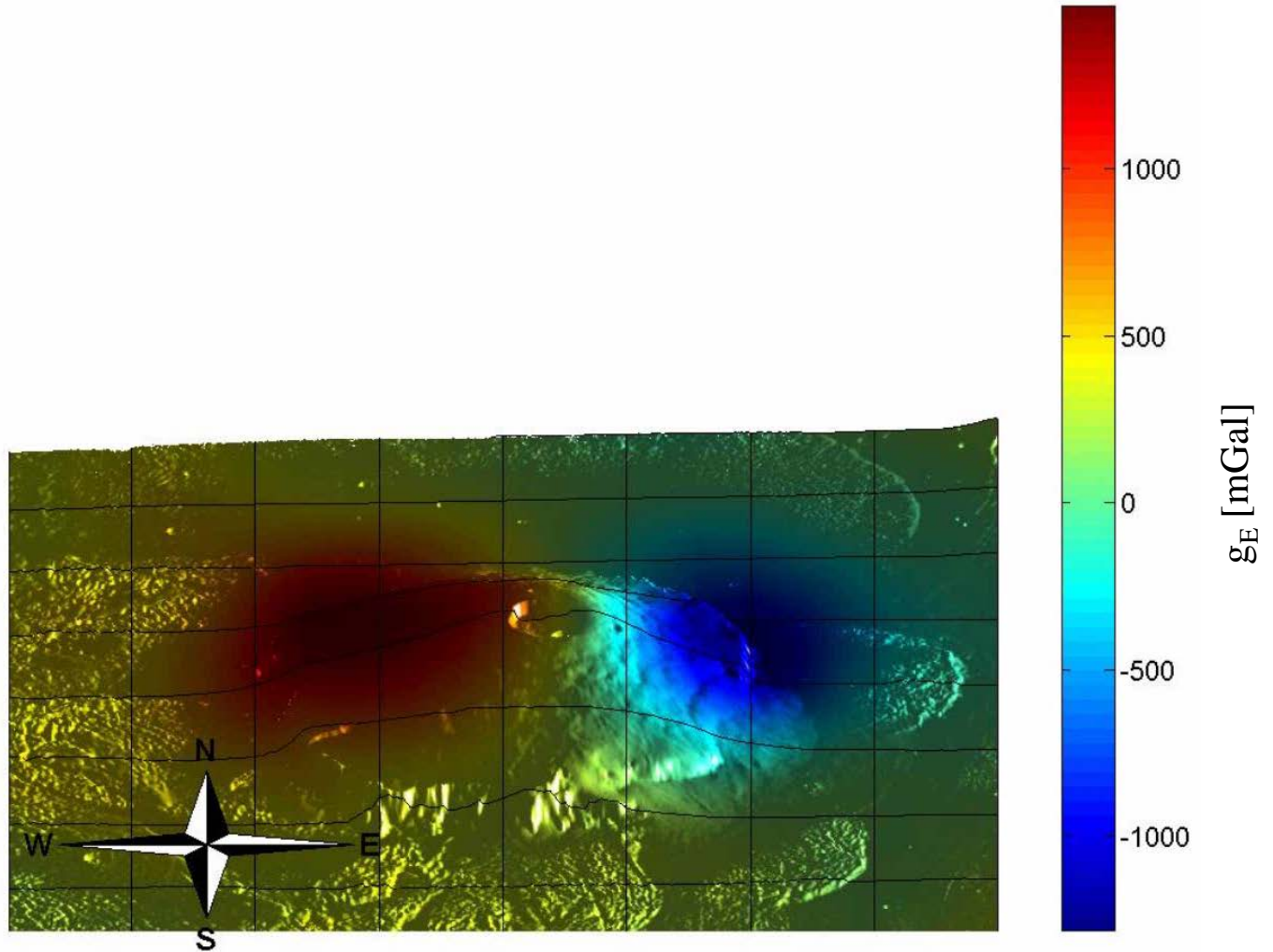
$$a_N = \frac{GM}{r^2} \sum_{l=2}^{\infty} \left(\frac{r_0}{r}\right)^l \sum_{m=0}^l \{P_{l,m+1}(\sin \delta) - m \tan \phi P_{lm}(\sin \delta)\} [C_{lm} \cos m\lambda + S_{lm} \sin m\lambda]$$

$$a_E = \frac{GM}{r^2 \cos \delta} \sum_{l=2}^{\infty} \left(\frac{r_0}{r}\right)^l \sum_{m=0}^l P_{lm}(\sin \delta) m [-C_{lm} \sin m\lambda + S_{lm} \cos m\lambda]$$

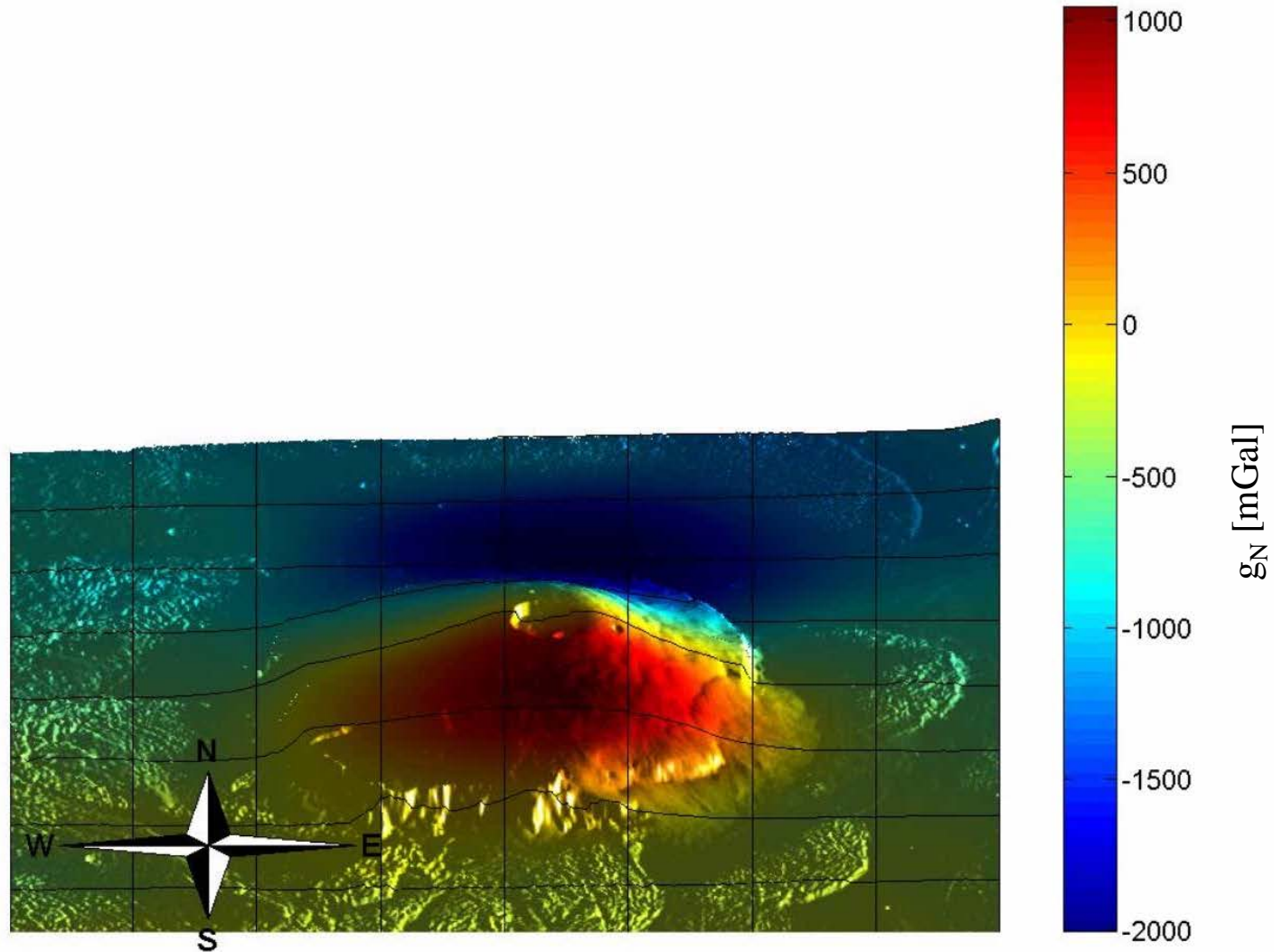
UBRZANJA GRAVITACIONOG POLJA



OLIMPUS MONS

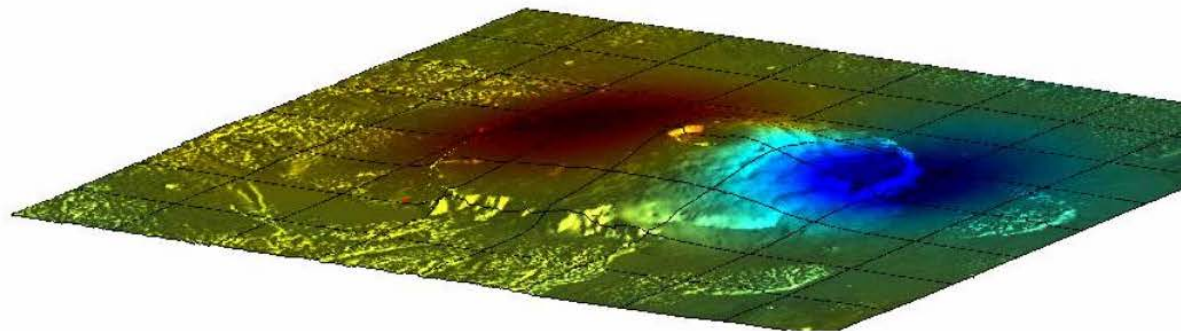


OLIMPIUS MONS



UTICAJ GRAVITACIONIH ANOMALIJA NA PUTANJU

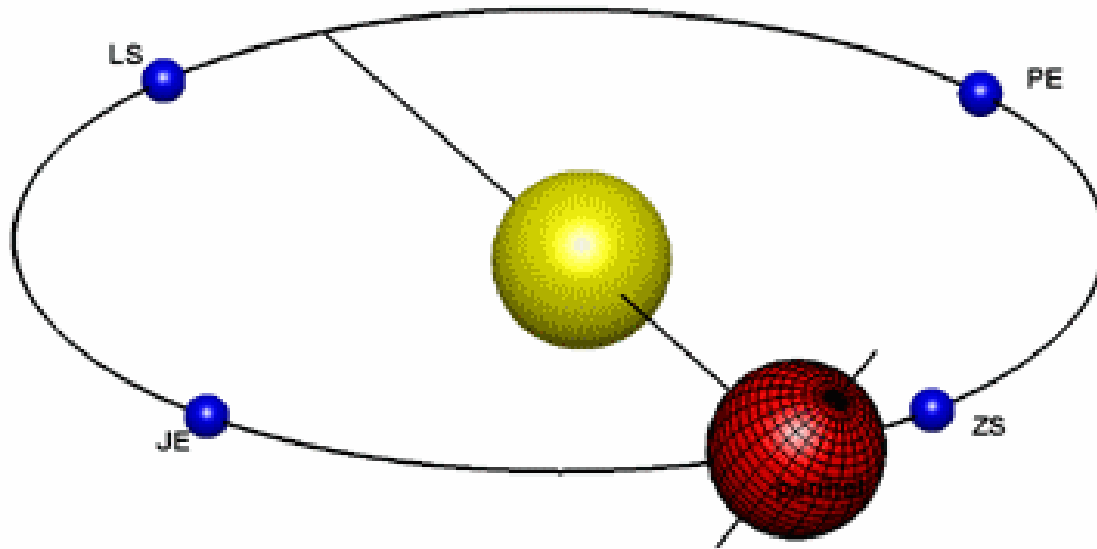
d=0 m



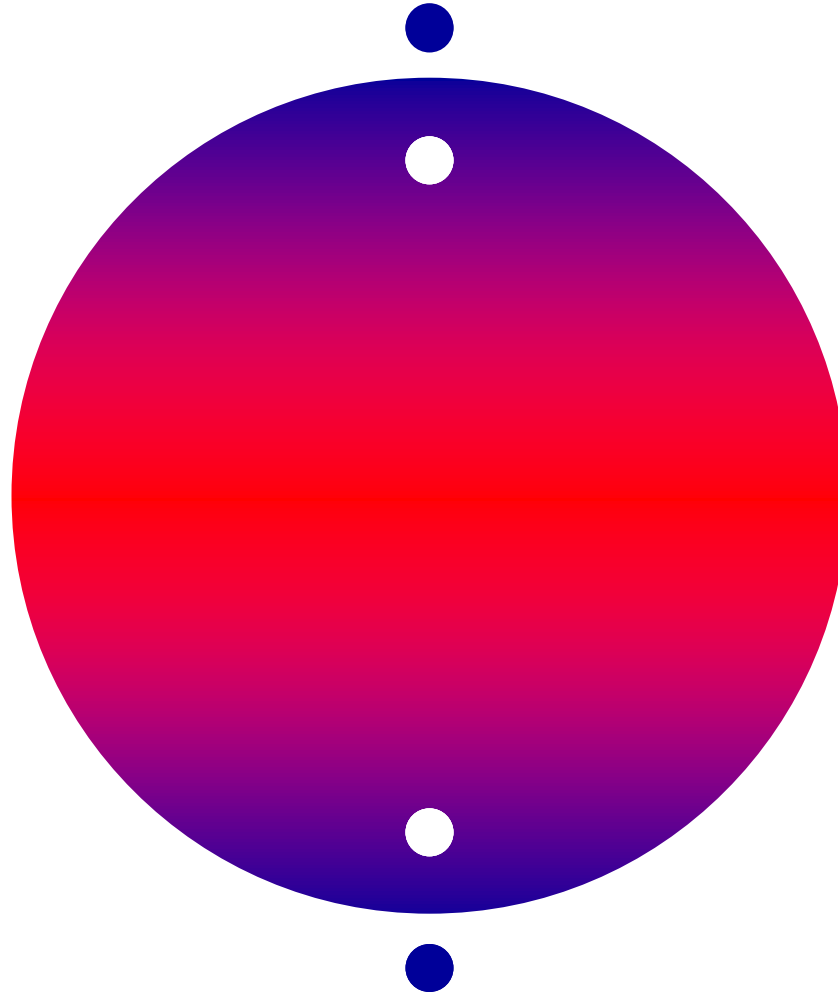
DANAŠNJA ATMOSFERA MARSA

	Zemlja	Mars	Odnos
p_0 [Pa]	101400	636	159
ρ_0 [kg/m ³]	1.217	0.02	60.8
H [km]	8.5	11.1	0.77

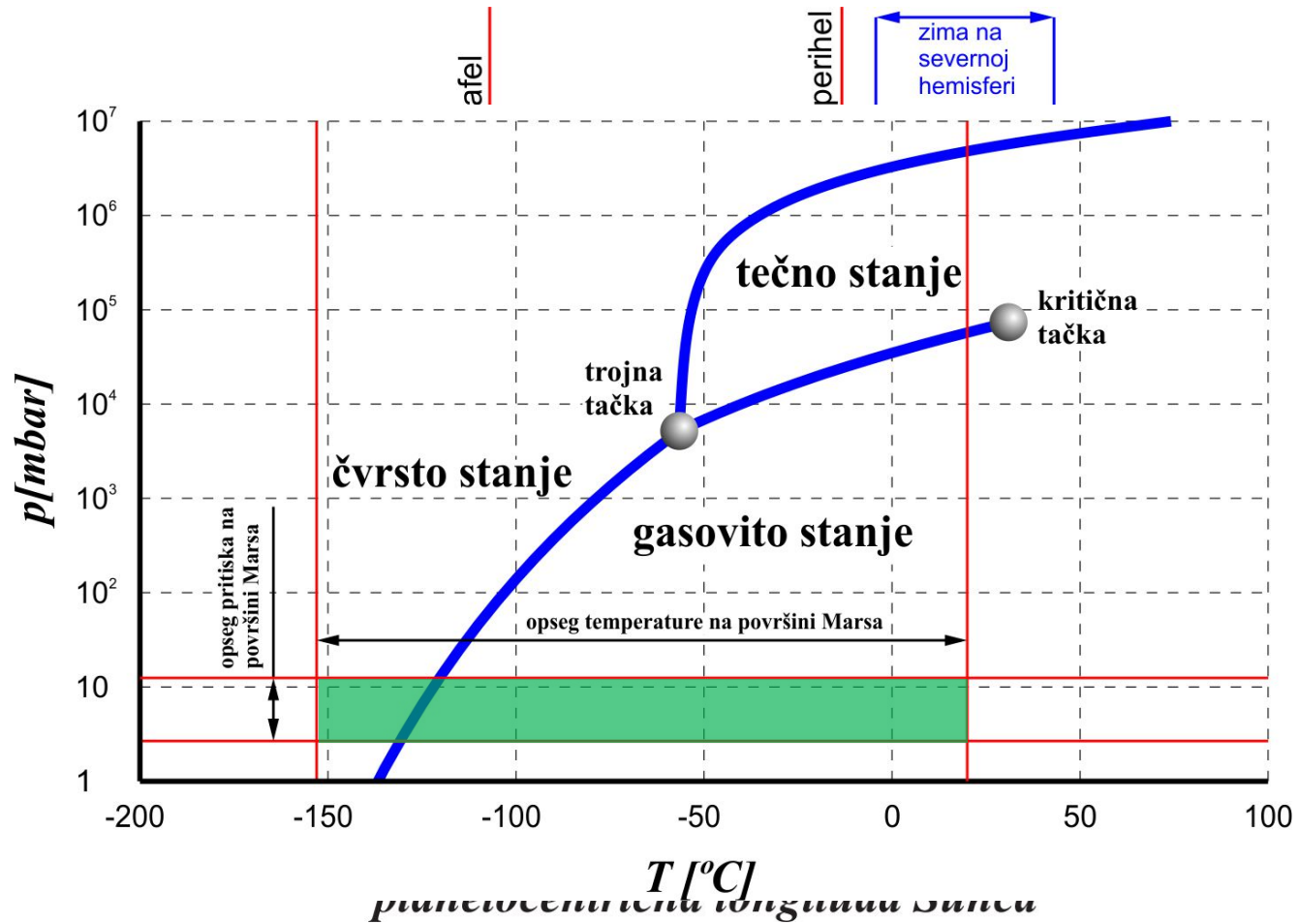
CIRKULACIJA ATMOSFERE



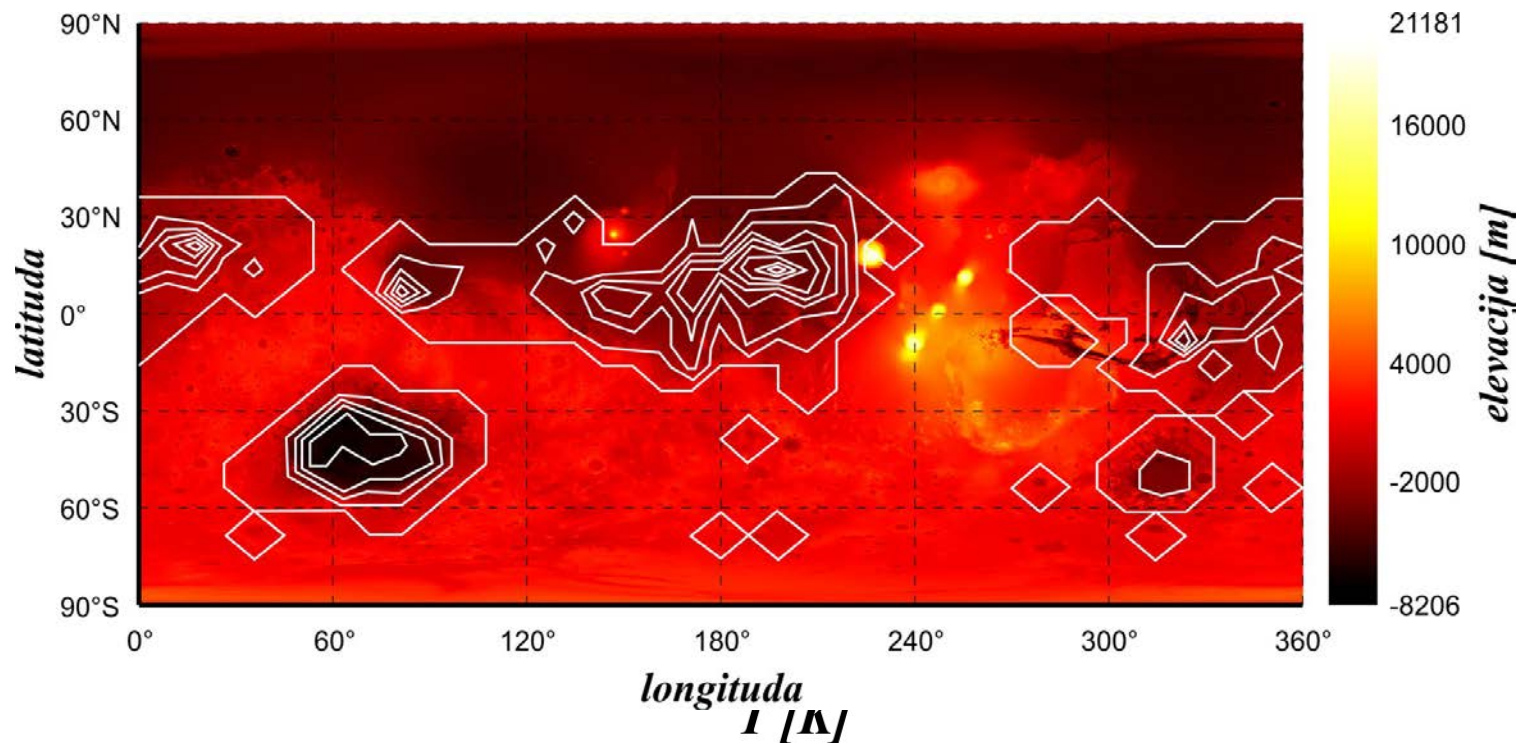
CIRKULACIJA ATMOSFERE



CO₂ CIKLUS



H₂O CIKLUS



Haberle et al., *Journal of Geophysical Research*, 2001.

MODELI CIRKULACIJE MARSOVE ATMOSFERE



$$\frac{D\rho}{Dt} + \rho \nabla \vec{v} = 0 \quad \text{Jednačina kontinuiteta}$$

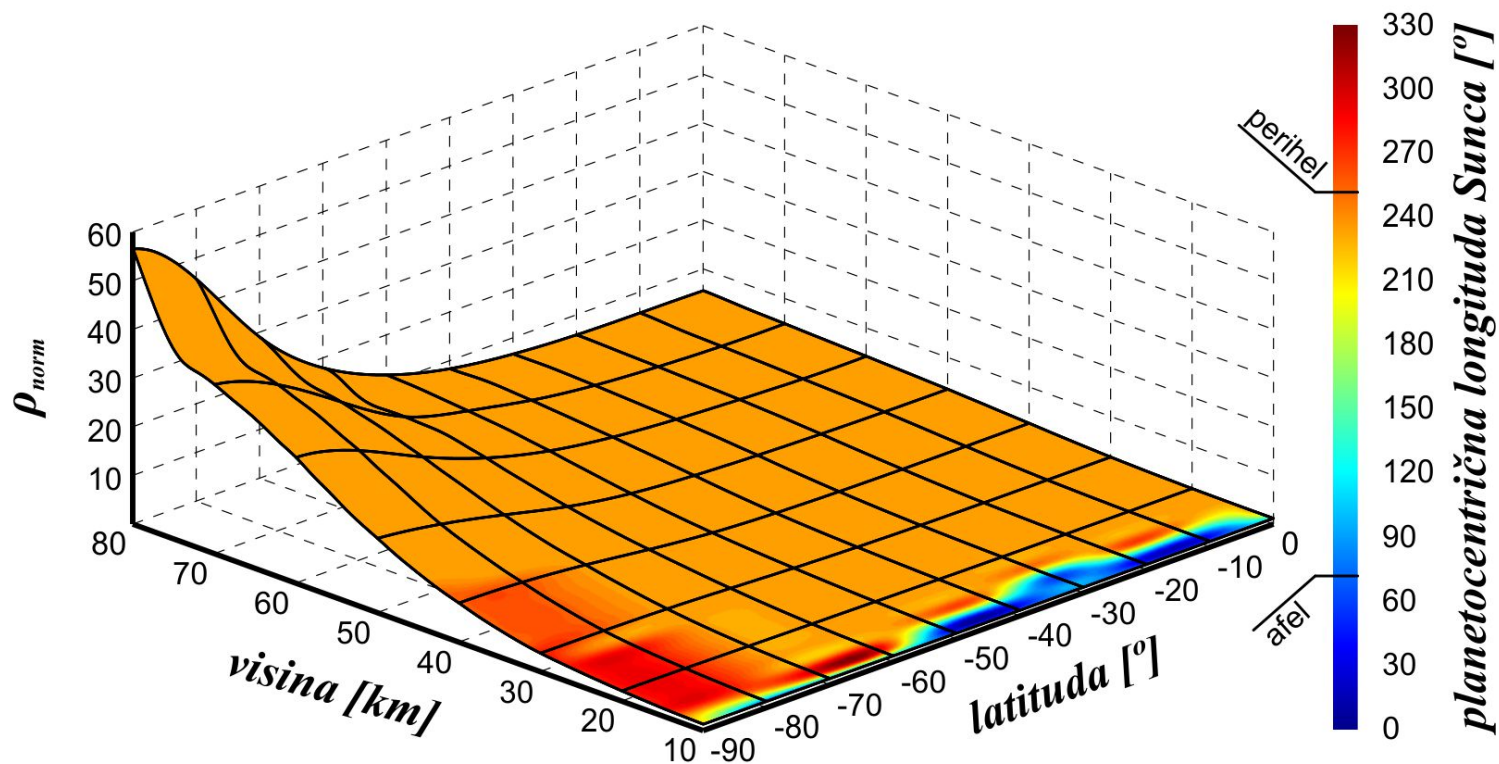
$$\frac{D\theta}{dt} = Q \quad \text{Jednačina energije}$$

Atmosferske baze podataka

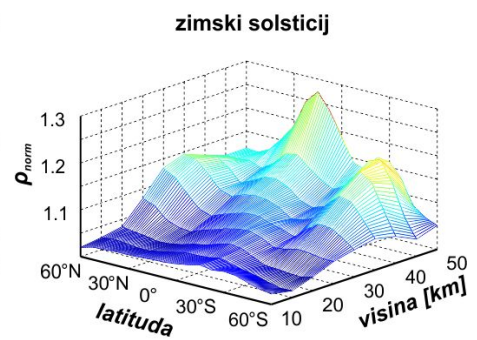
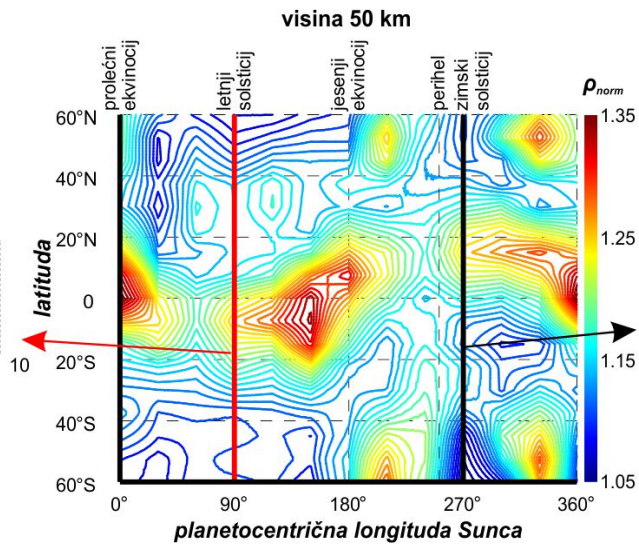
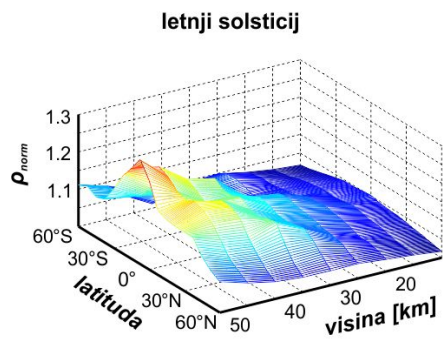
$$\frac{D\vec{v}}{Dt} = -2\Omega \times \vec{v} - \Omega \times (\Omega \times r) - \vec{g} - \frac{1}{\rho} \nabla p + \frac{\eta_d}{\rho} \nabla^2 \vec{v} \quad \text{N-S jednačina}$$

SEZONSKI CIKLUS

Severna hemisfera



DNEVNI CIKLUS



AERODINAMIKA

$$F_x = \frac{1}{2} \rho v^2 S C_x$$

PARADOKS TUPOG TELA

CO₂, N₂, Ar

Ar, C, N, O, C₂, N₂, O₂, CN,
CO, NO, CO₂, NCO, Ar⁺,
C⁺, N⁺, O⁺, C₂⁺, N₂⁺, O₂⁺,
CN⁺, CO⁺, NO⁺

v_1
 $\xrightarrow{M_1 > 1}$

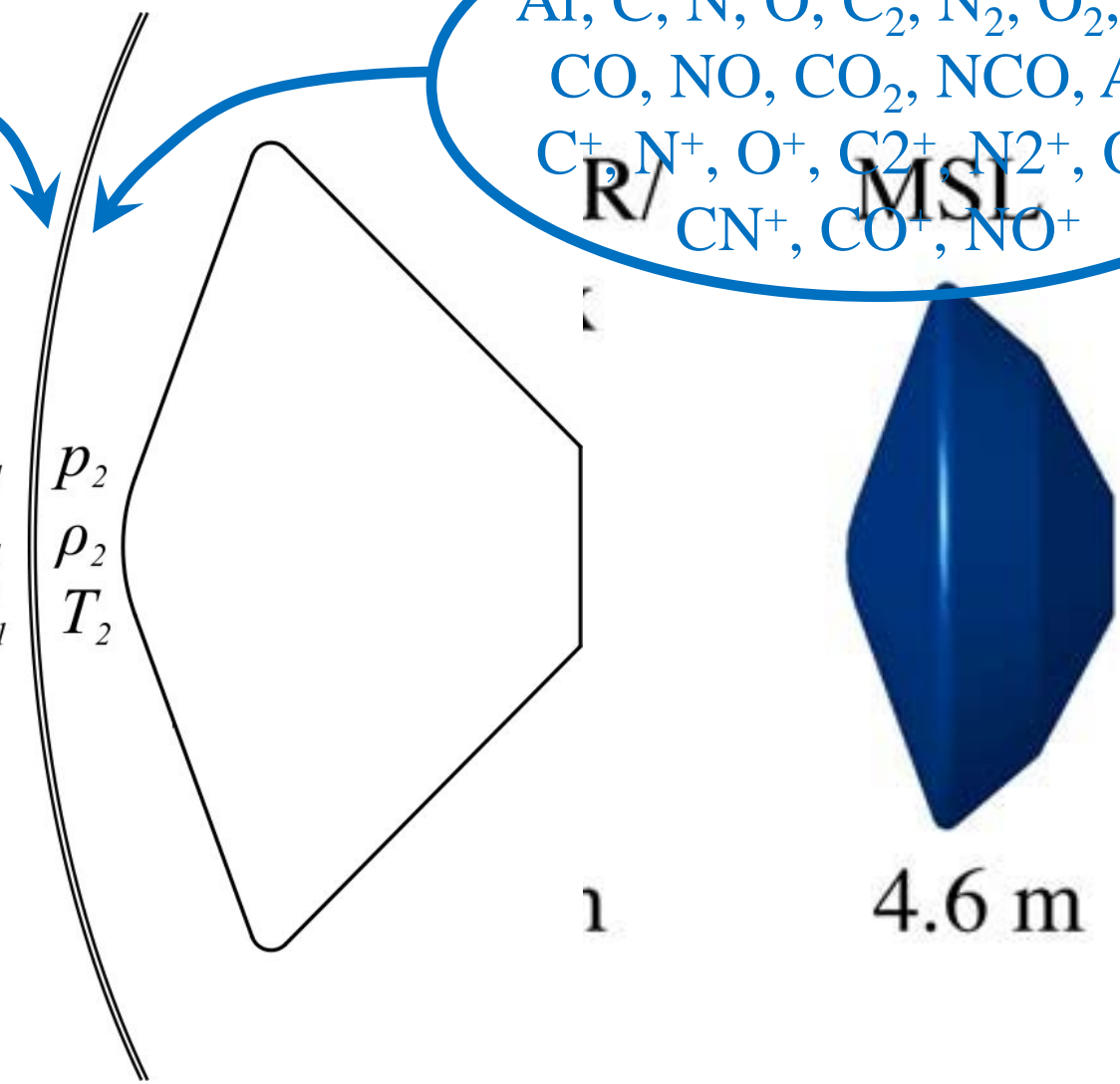
p_1 p_2
 ρ_1 ρ_2
 T_1 T_2

R/

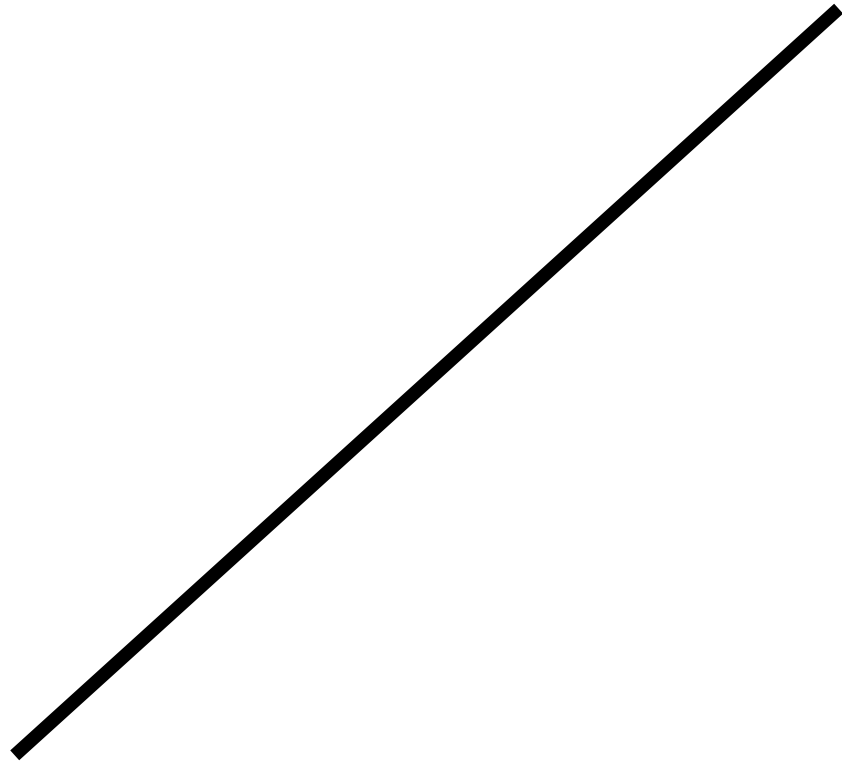
MSL

1

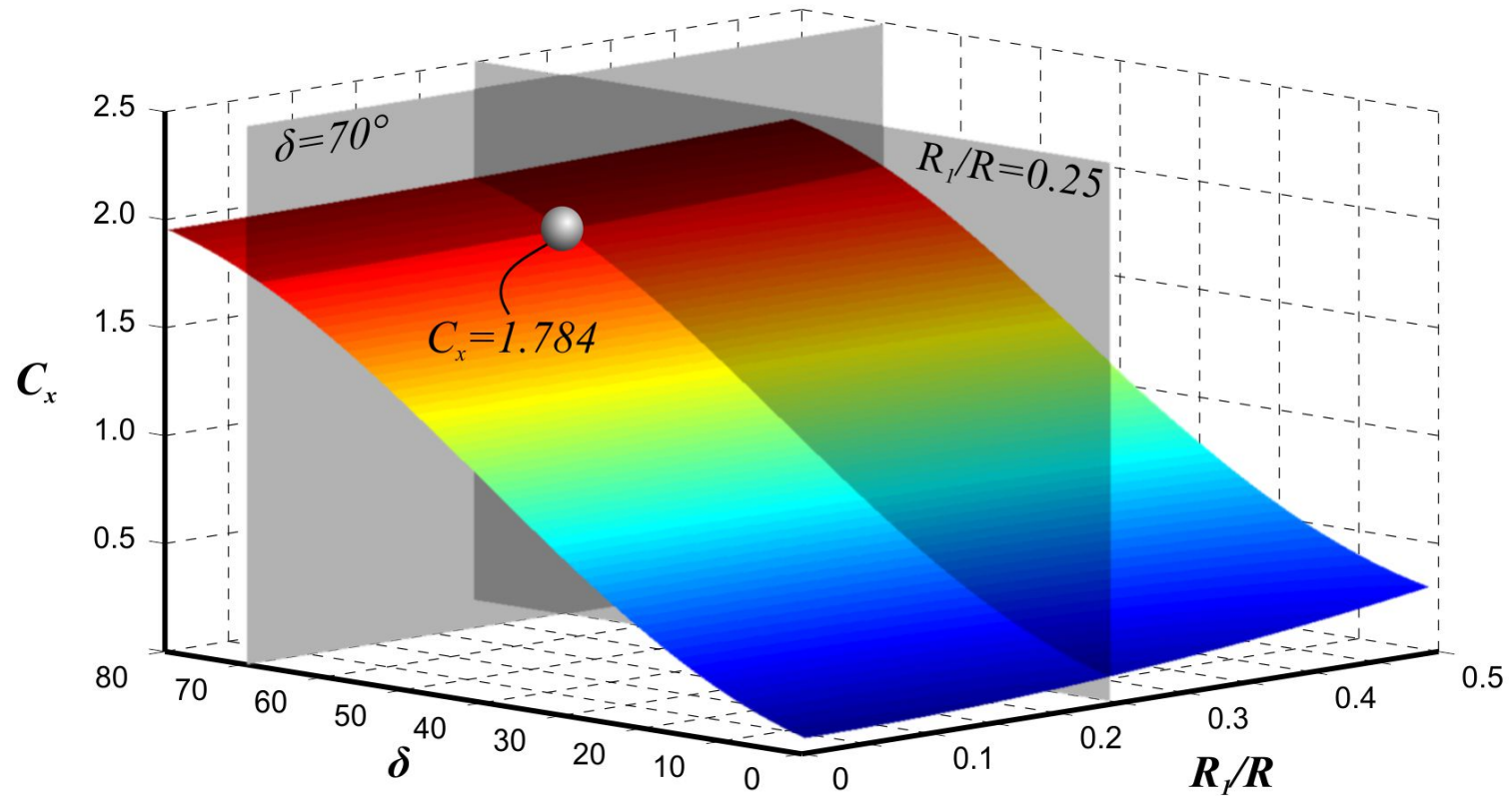
4.6 m



NJUTNOVA TEORIJA



NJUTNOVA TEORIJA

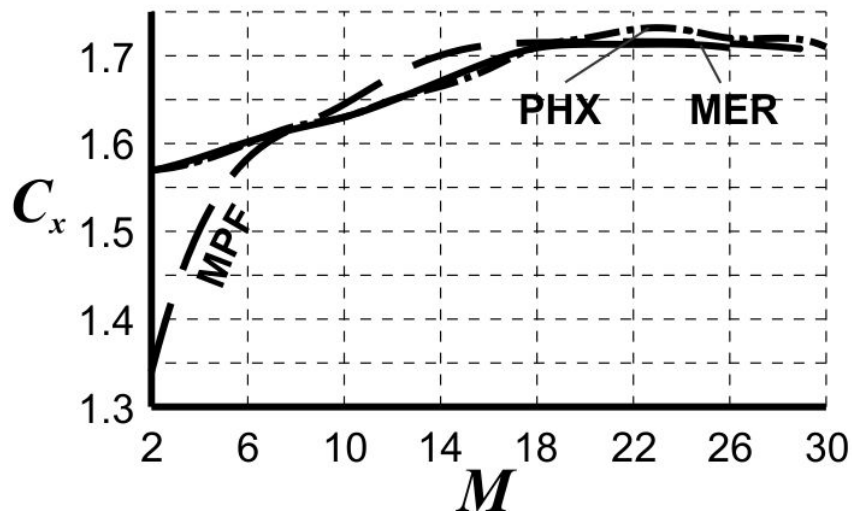


$$C_x = \frac{1}{S} \iint_A C_p dA = \frac{2}{R^2 \pi} \iint_A \sin^2 \theta dA$$

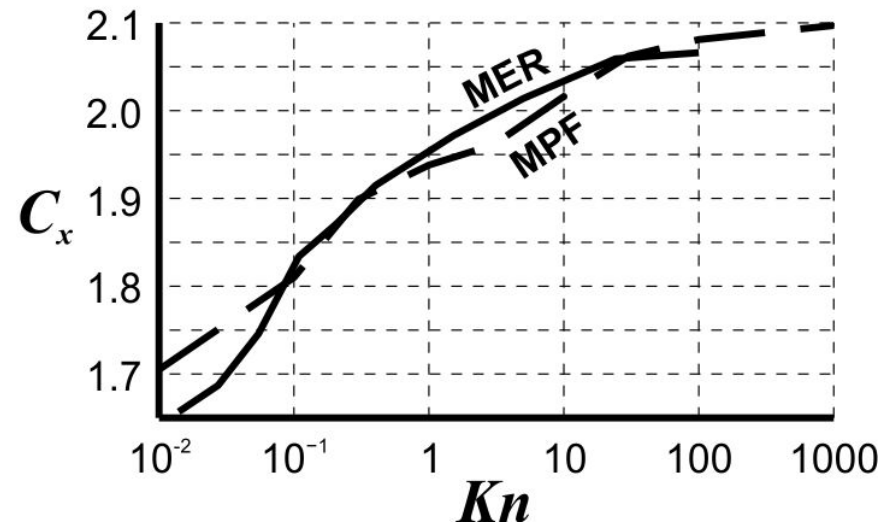


AERODINAMIČKI MODEL

Kontinualni režim



prelazni i slobodno-molekularni režim



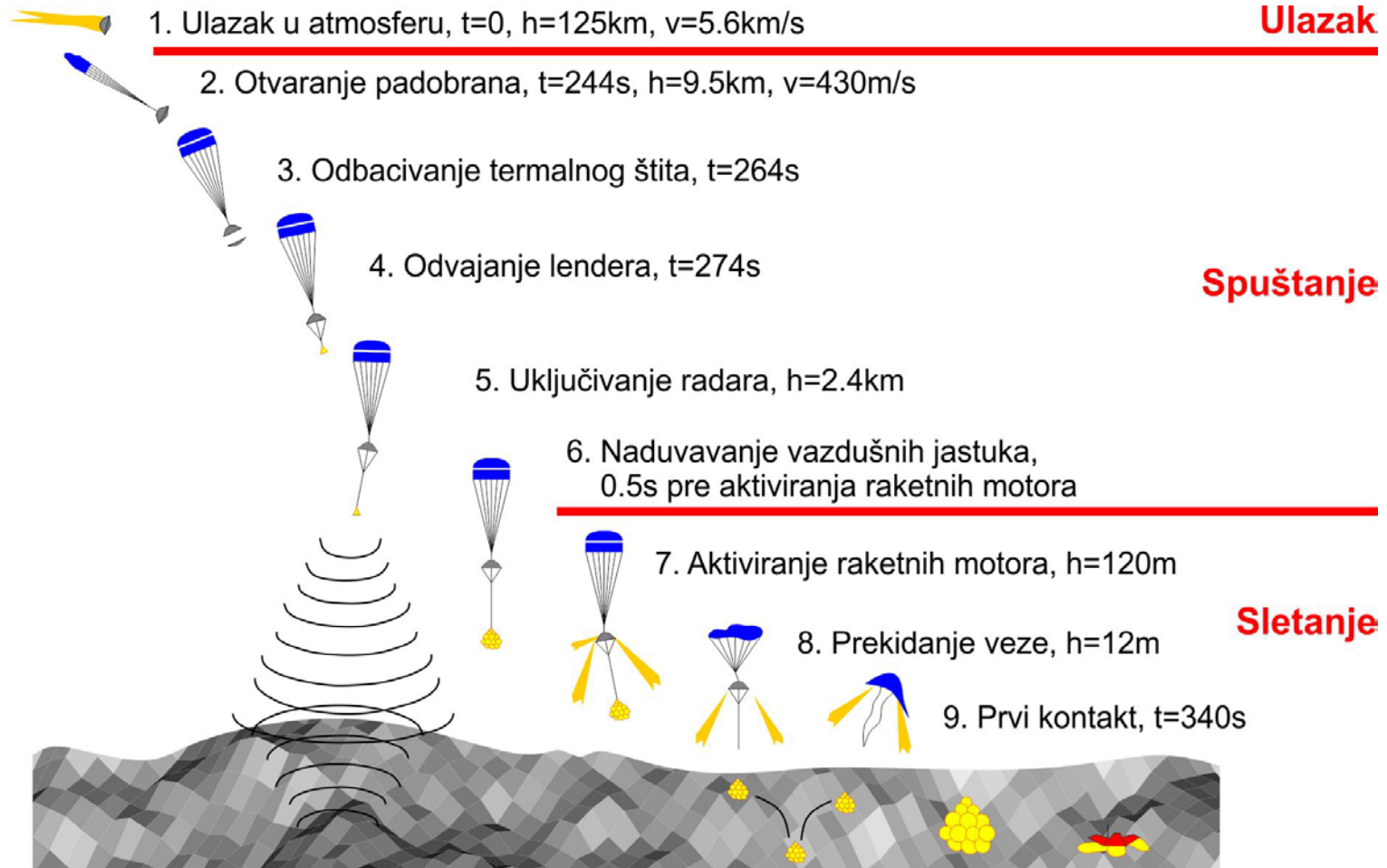
Gnoffo et al., *Journal of Spacecraft and Rockets*, 1996.

Moss et al., *AIAA paper 98-0298*, 1998.

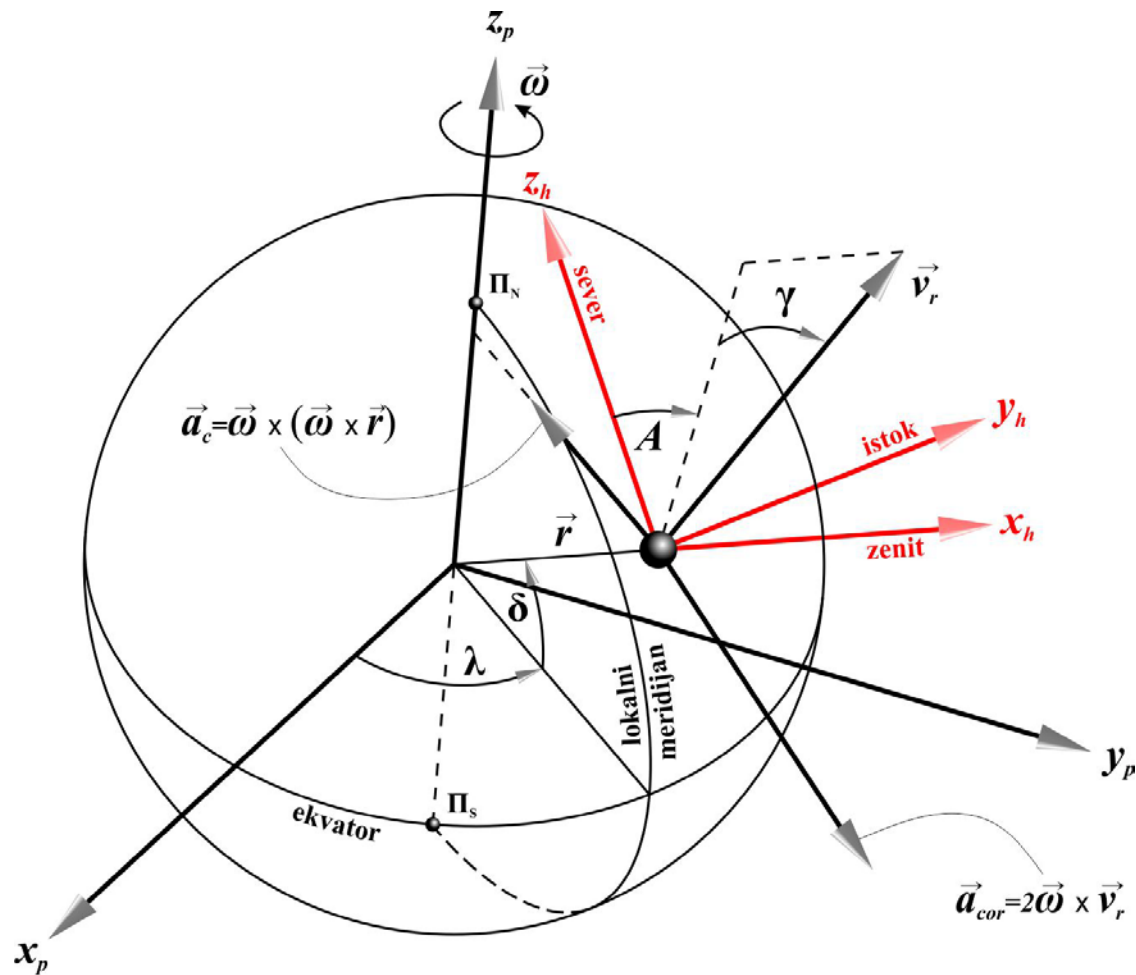
Edquist & Desai, *Journal of Spacecraft and Rockets*, 2011.

Schoenenberger & Cheatwood, *AIAA paper 2005-0056*, 2005.

DINAMIKA ULASKA U ATMOSFERU, SPUŠTANJA I SLETANJA NA POVRŠINU MARSA

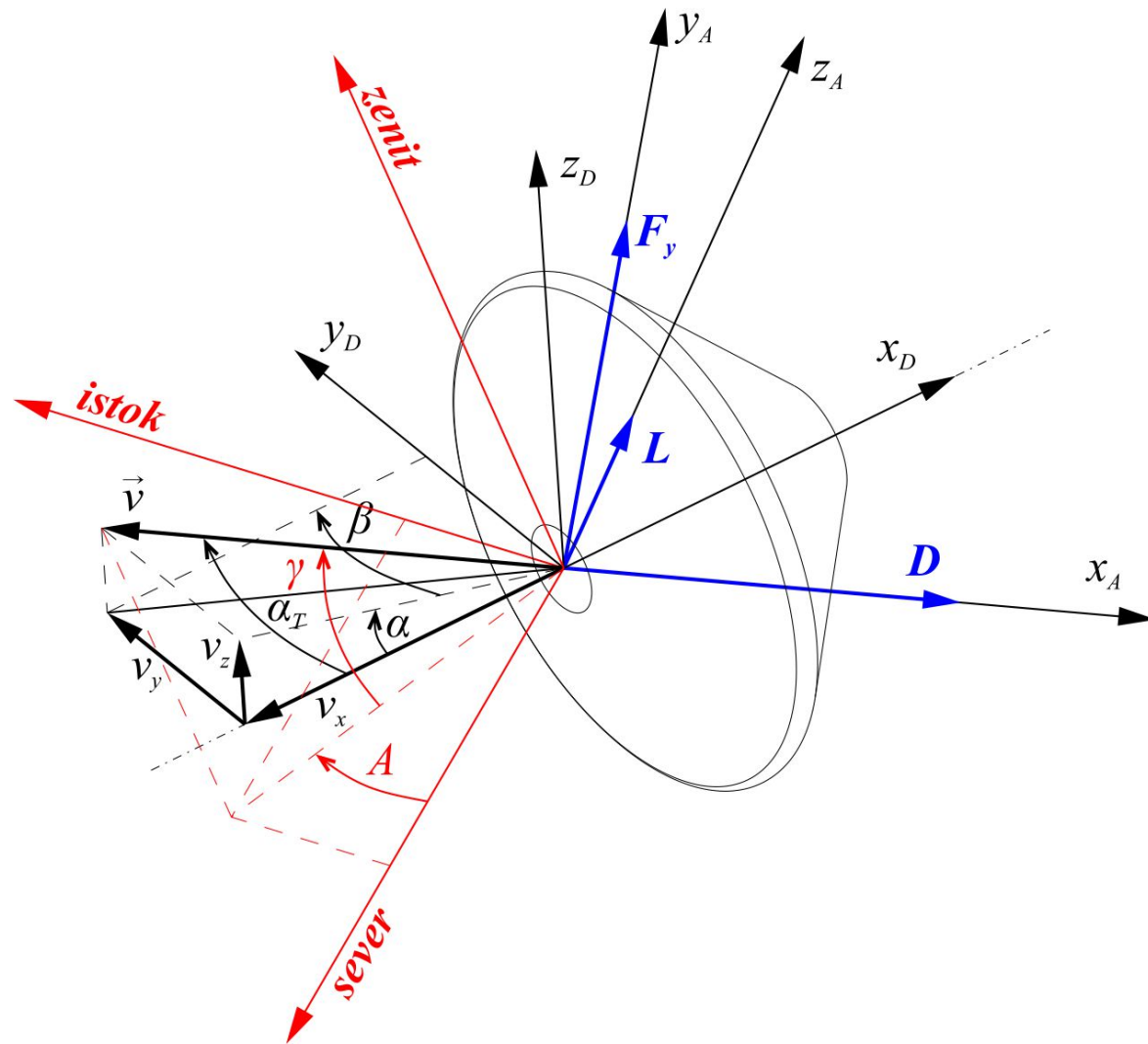


KOORDINATNI SISTEMI



$$\vec{a} = \vec{a}_r + 2\vec{\omega} \times \vec{v}_r + \vec{\omega} \times (\vec{\omega} \times \vec{r})$$

KOORDINATNI SISTEMI



JEDNAČINE KRETANJA

$$\dot{v} = -\frac{1}{m} F_x - g_C \sin \gamma + \cancel{g_N \cos \gamma \cos A}$$

$$\cancel{g_E \cos \gamma \sin A - \omega^2 r \cos \delta (\cos \gamma \cos A \sin \delta - \sin \gamma \cos \delta)}$$

$$\dot{v} = -\frac{1}{2m} \rho v^2 C_x - g \sin \gamma$$

$$\dot{A} = \frac{v}{r} \cos \gamma \sin A \tan \delta + \frac{1}{v \cos \gamma} (\cancel{\omega^2 r \sin A \sin \delta \cos \delta - g_N \sin A} + \cancel{g_E \cos A})$$

$$-2\omega (\cancel{\tan \gamma \cos A \sin \delta} + \cos \delta)$$

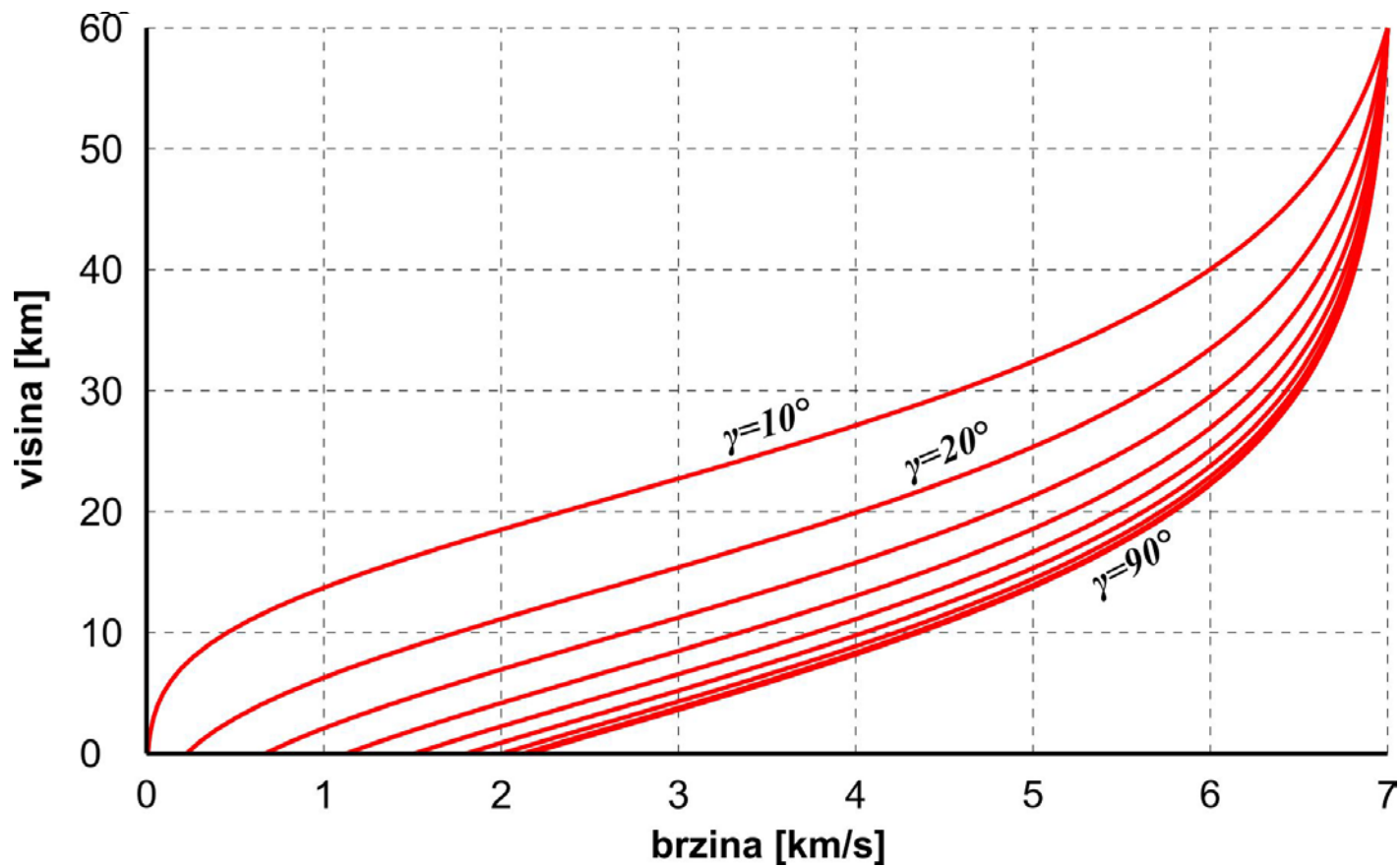
$$\dot{\gamma} = \cos \gamma \left(\frac{v}{r} - \frac{g}{v} \right)$$

$$\dot{\gamma} = \frac{v}{r} \cos \gamma$$

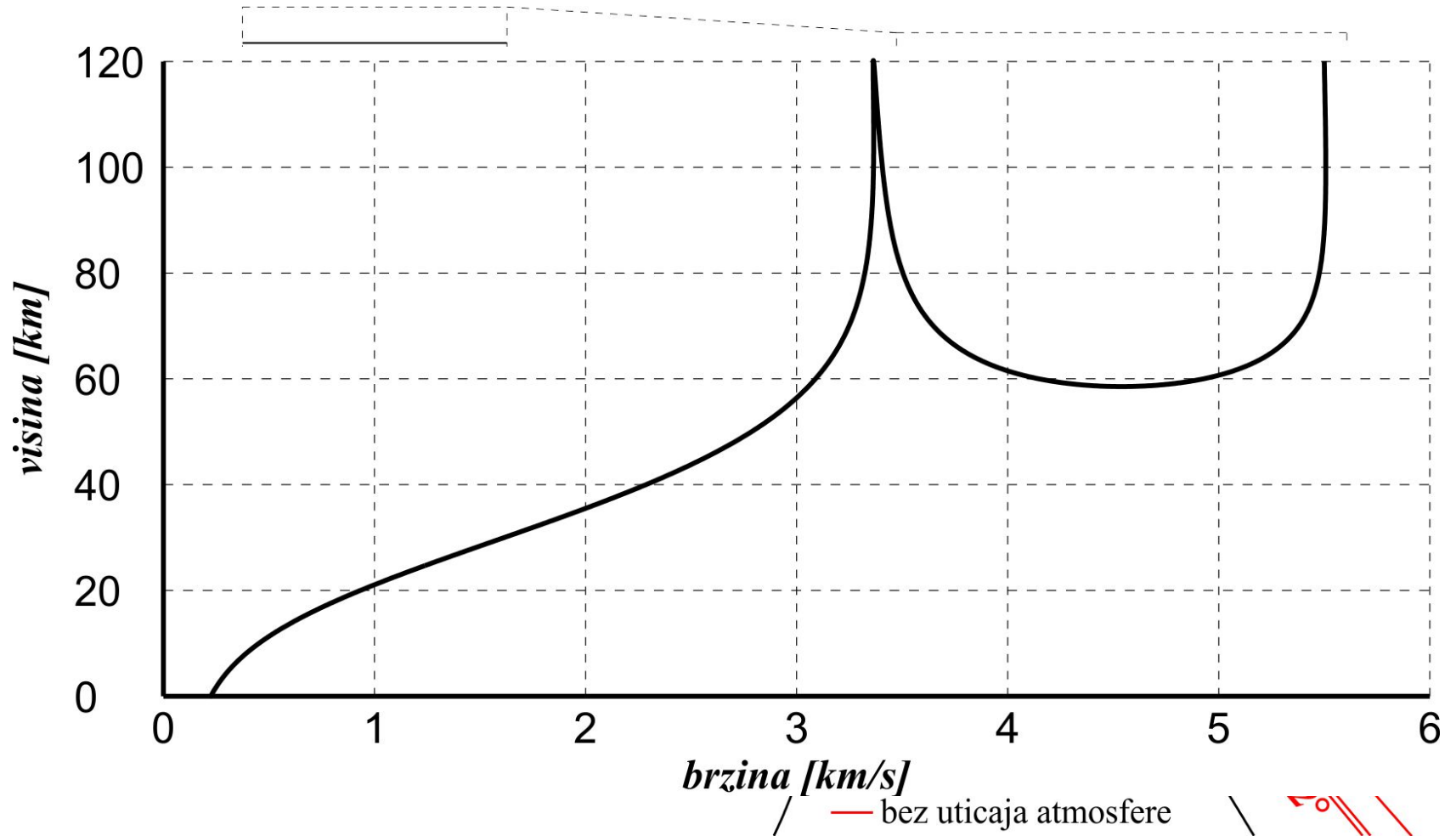
$$-\frac{1}{v} [g_C \cos \gamma + \cancel{g_N \sin \gamma \cos A} + \cancel{g_E \sin \gamma \sin A} - \cancel{\omega^2 r \cos \delta (\sin \gamma \cos A \sin \delta + \cos \gamma \cos \delta)}]$$

$$+2\omega \cancel{\sin A \cos \delta}$$

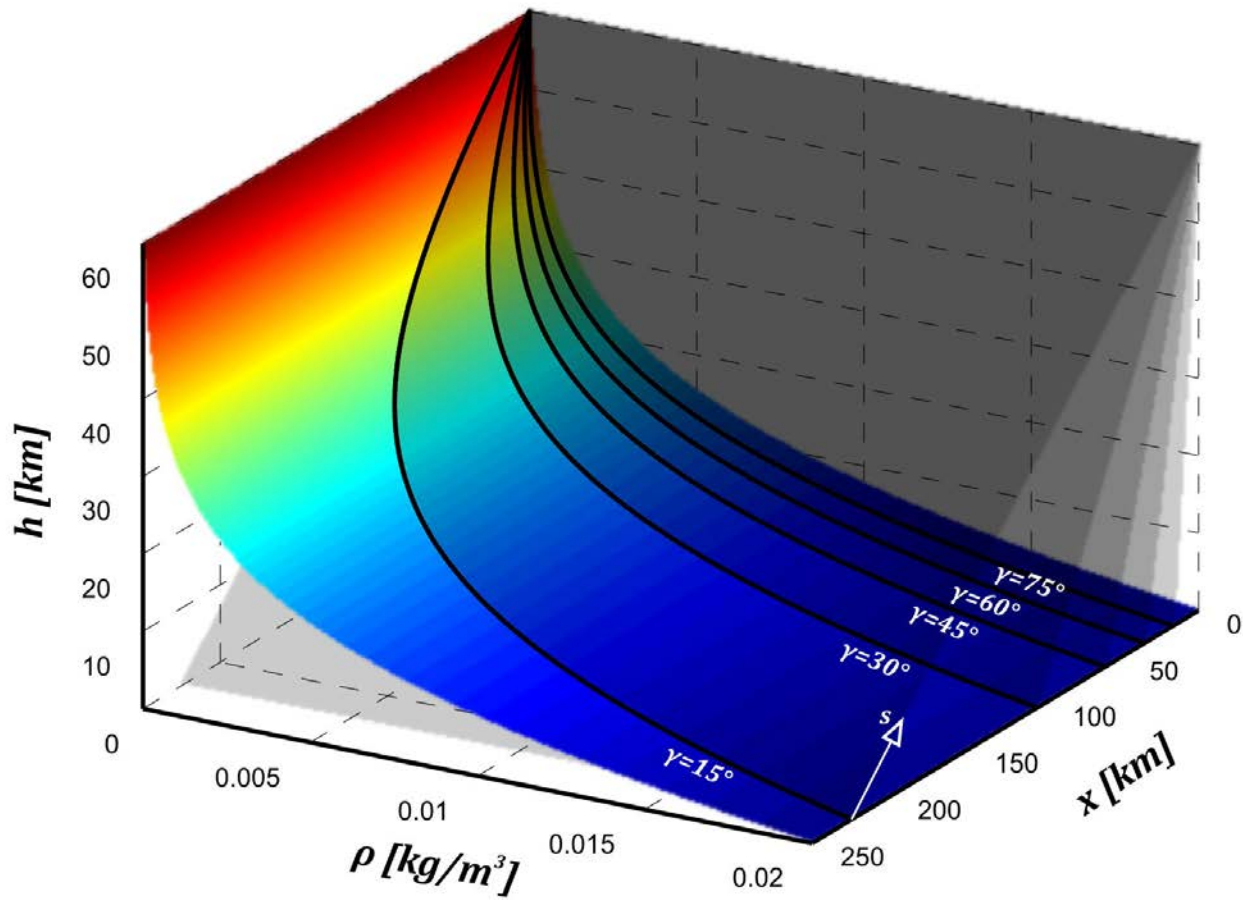
ANALITIČKO REŠENJE



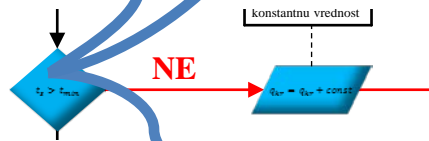
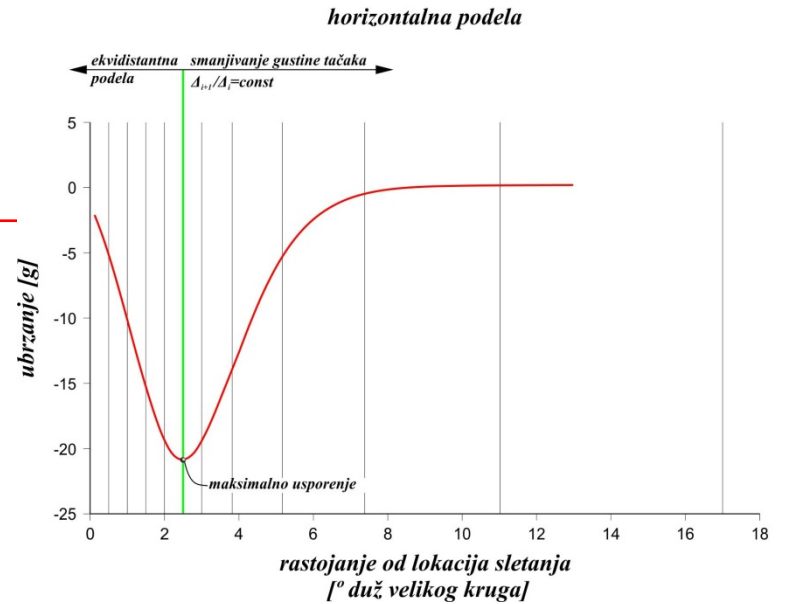
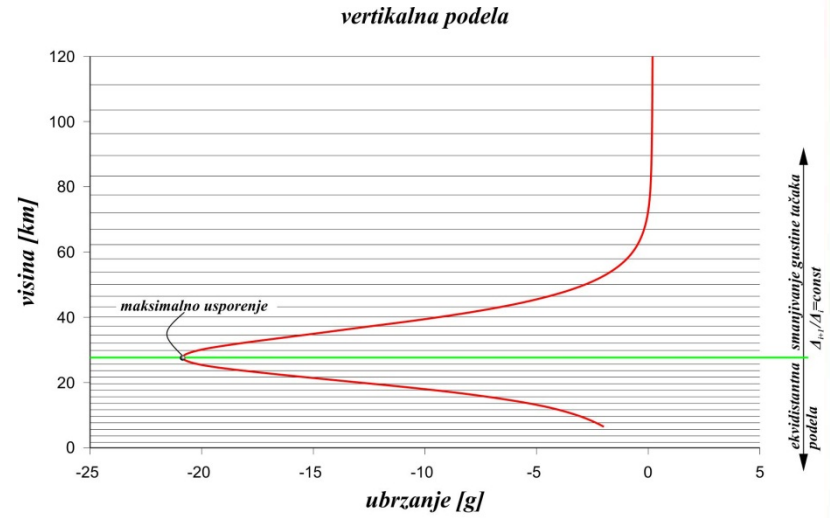
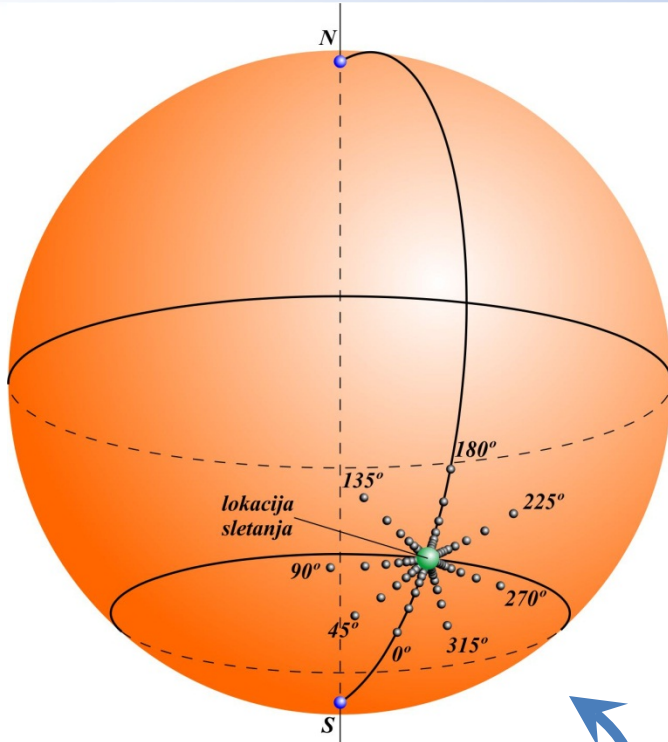
GRANIČNA PUTANJA



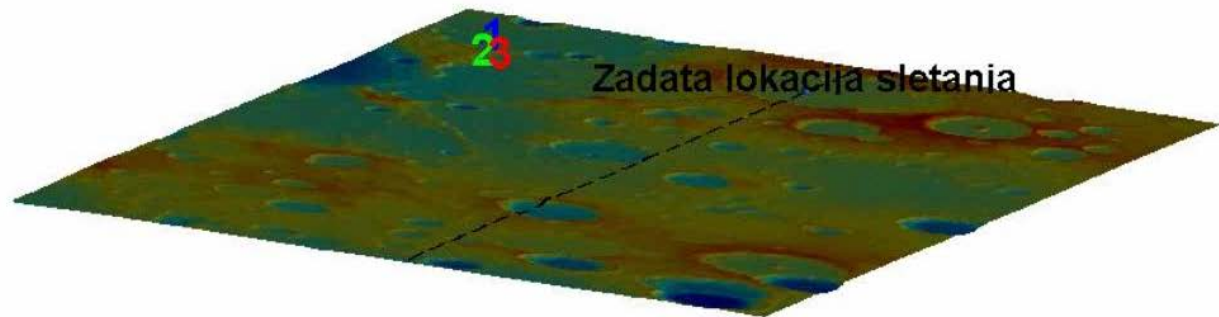
„PRIVIDNA“ ATMOSFERA



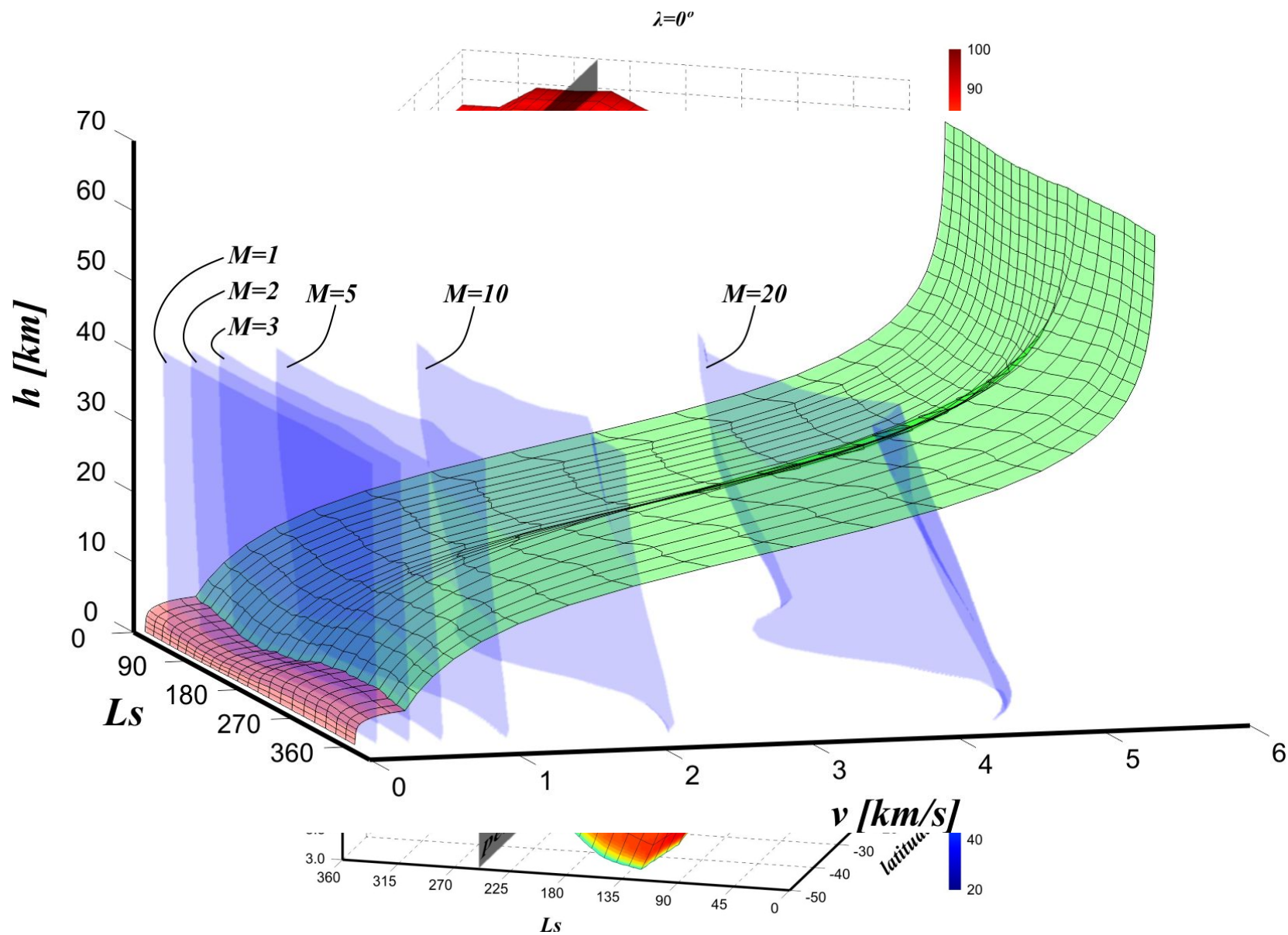
ODREĐIVANJE OPTIMALNE PUTANJE



KONVERGENCIJA PUTANJE



UTICAJ SEZONSKOG CIKLUSA

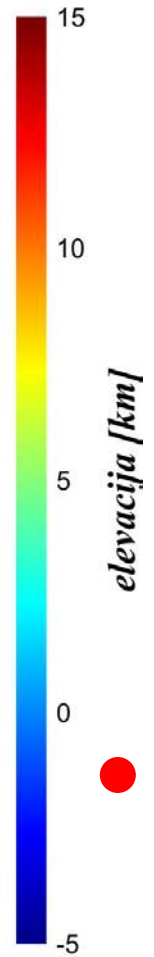
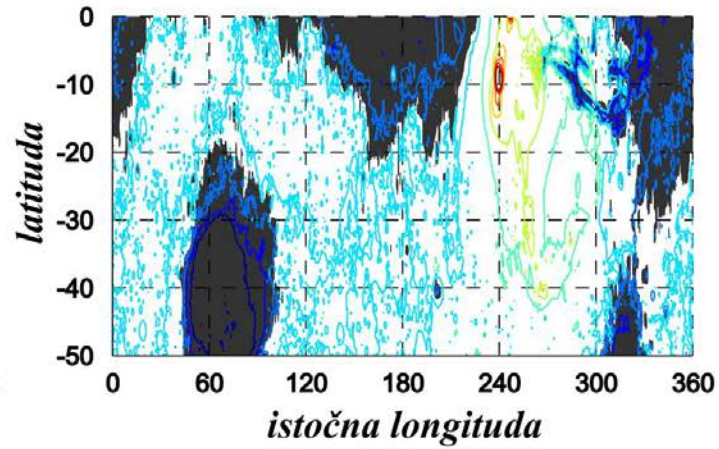
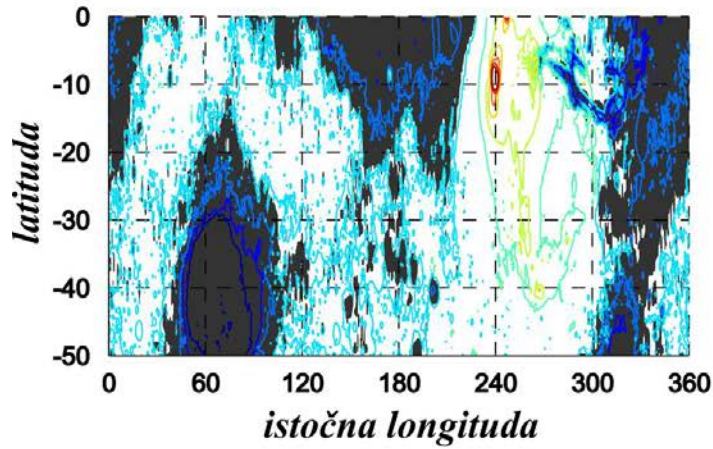


PRISTUPAČNE LOKACIJE

$\lambda=0^\circ$

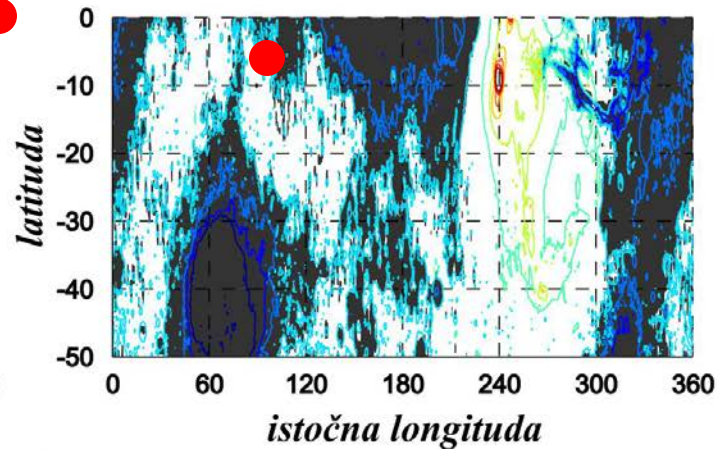
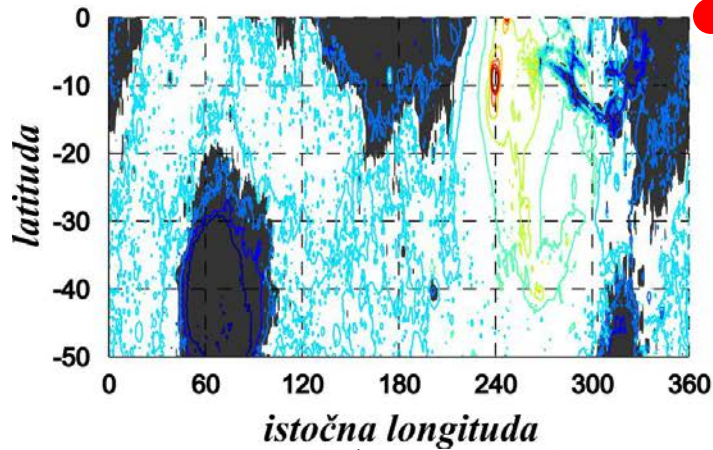
$Ls=0^\circ$

$Ls=90^\circ$



$Ls=180^\circ$

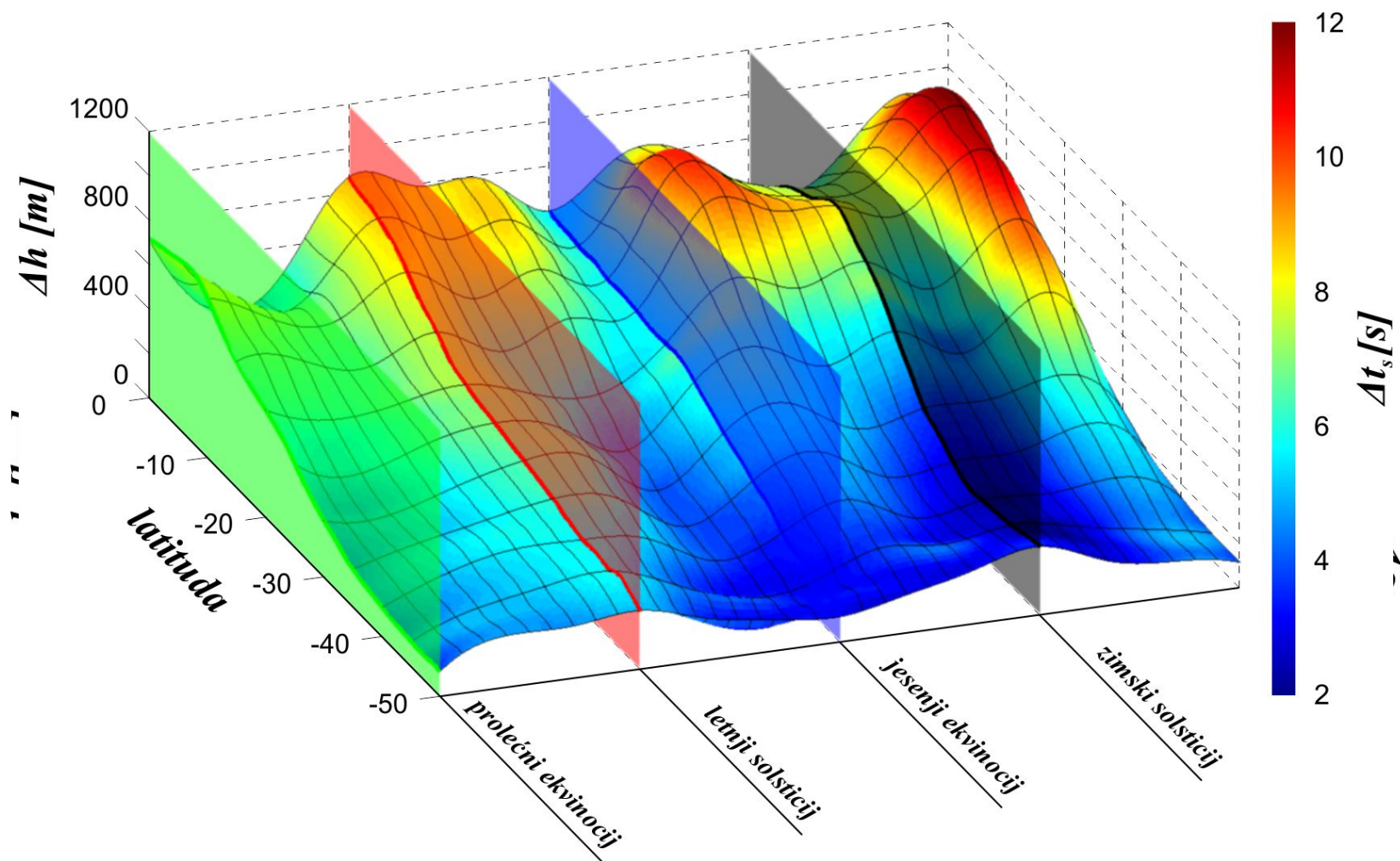
$Ls=270^\circ$



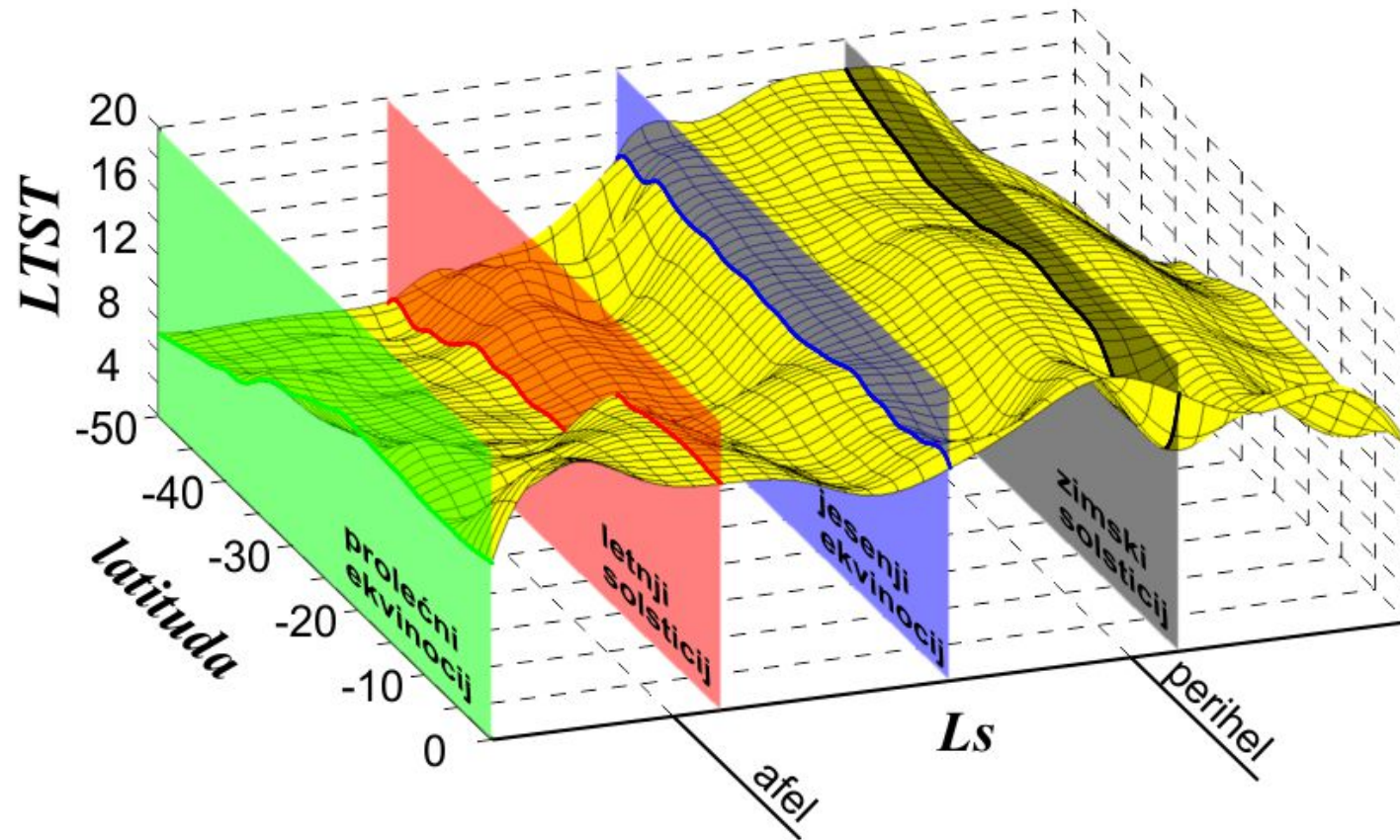
-50 0 90 180 270 360

Ls

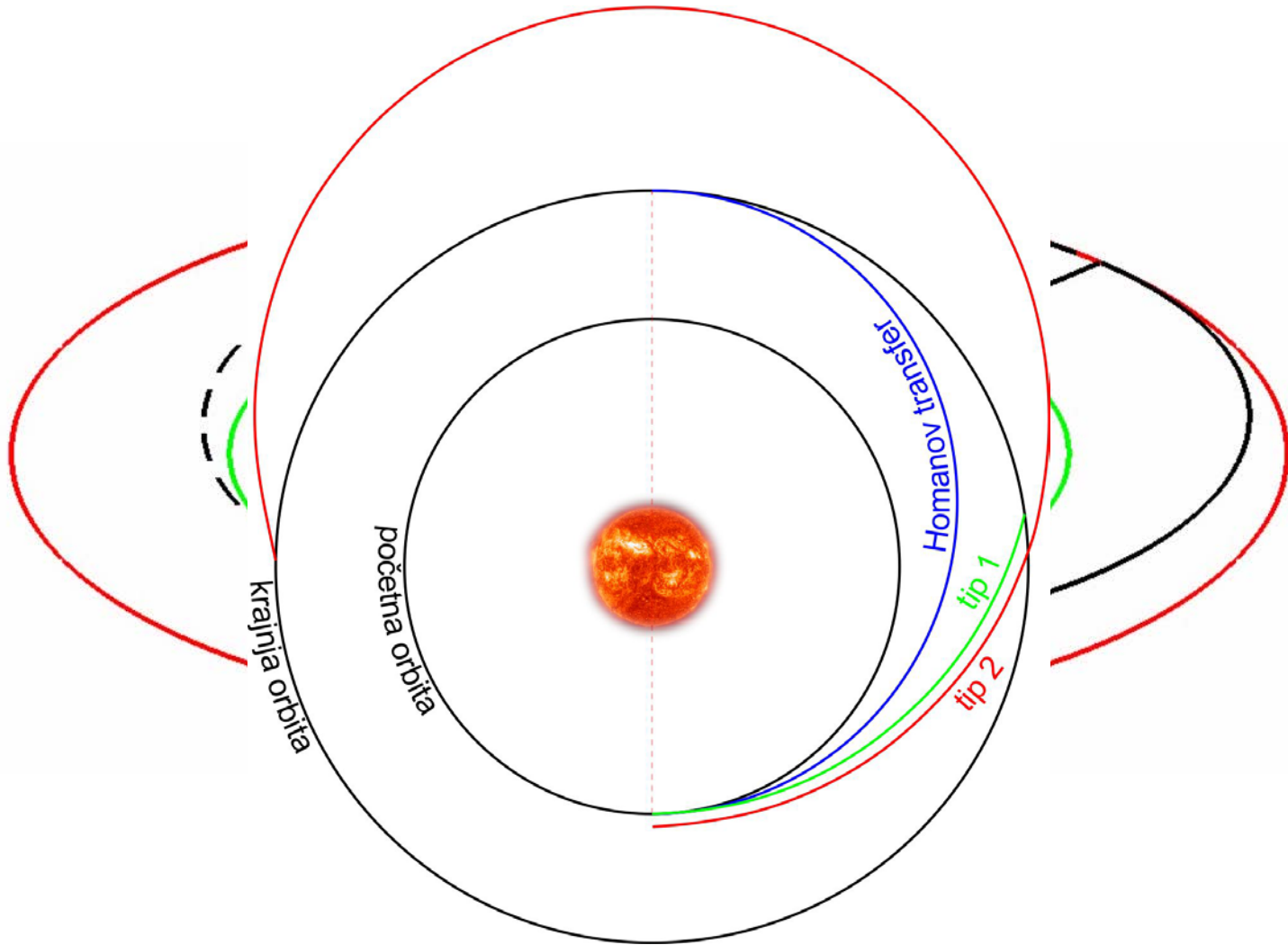
UTICAJ DNEVNOG CIKLUSA



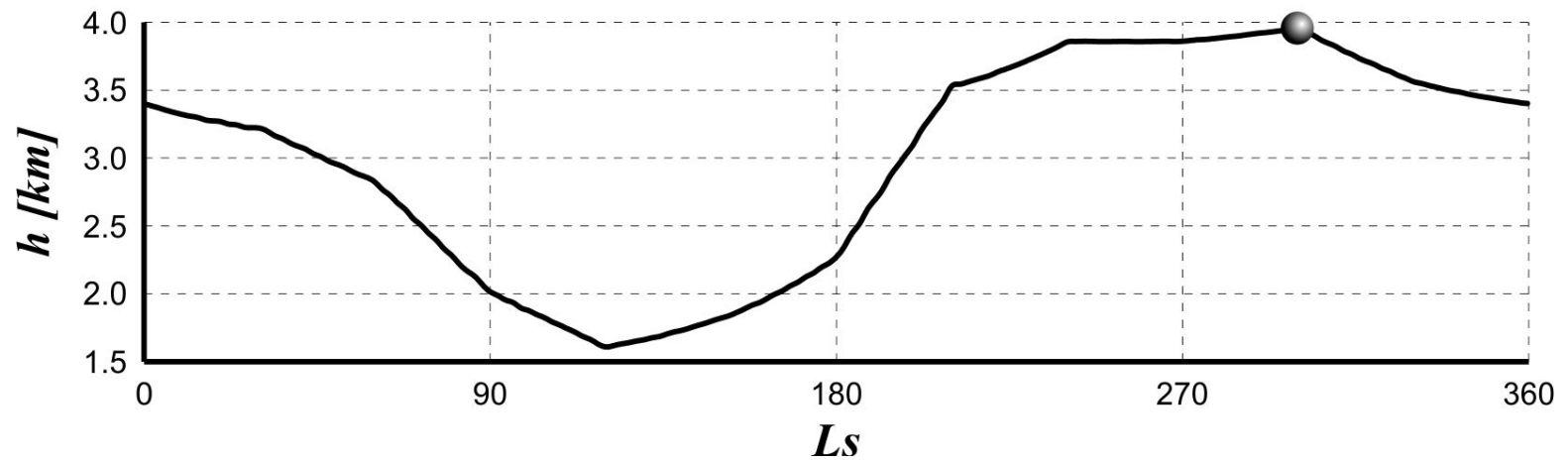
OPTIMALNO LPSV



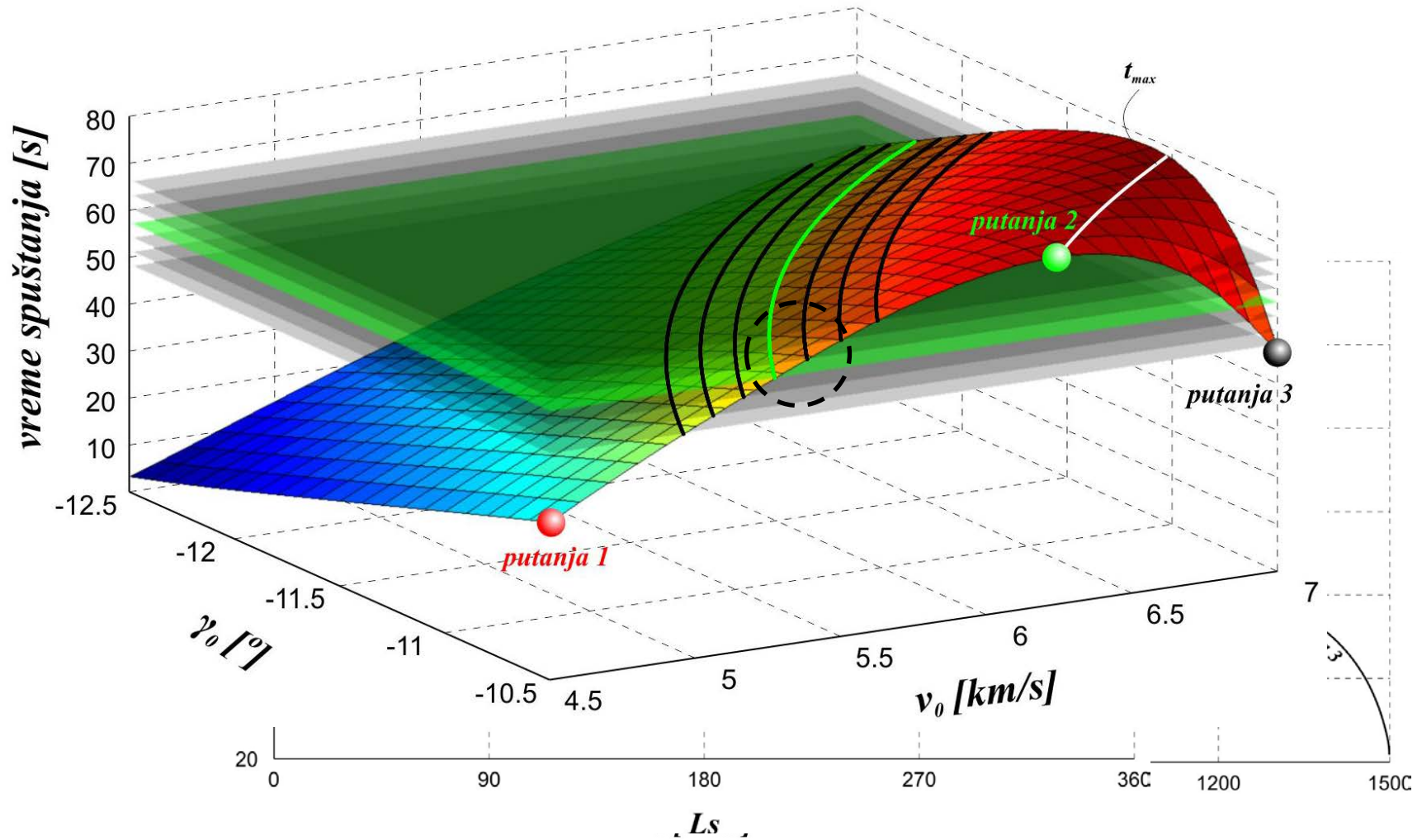
MEĐUPLANETARNI TRANSFER



OPTIMIZACIJA POČETNIH USLOVA



OPTIMIZACIJA POČETNIH USLOVA



SAŽETAK

- Zbog sezonskih varijacija atmosferskih parametara, pre svega gustine, koje su dominantno uzrokovane CO₂ ciklusom u Marsovoj atmosferi, pristupačne elevacije sletanja se menjaju za oko 2.5 *km* u toku godine.
- Optimalna sezona za sletanje na južne visoravni je u okolni perihela orbite, od kasnog proleća do rane jeseni.
- Uticaj sezonskih varijacija atmosferskih parametara raste sa udaljavanjem od ekvatora.
- Pristupačne lokacije sletanja pojavljuju se u visokim oblastima, na visoravnima Sirenum, Simeria i Noakis, kada je Mars u okolini perihela.
- Optimizacijom početnih uslova (v_0, γ_0) može se postići sletanje na lokacije sa velikom elevacijom bez unapređenja postojeće tehnologije.
- Početni uslovi (v_0, γ_0) imaju paraboličku zavisnost na izohronama spuštanja.
- Uticaj dnevnih varijacija atmosferskih parametara je višestruko manji od sezonskih i, suprotno od njih, opada sa udaljavanjem od ekvatora.
- Optimalno vreme sletanja je uvek tokom obdanice.

HVALA NA PAŽNJI

